

Pontecorvo, Neutrinos, MINOS

Pontecorvo Prize Talk
JINR, Dubna
February 16, 2012

Stanley Wojcicki
Stanford University



Bruno Pontecorvo

1913-1993



Bruno Pontecorvo

1913-1993



- Solar neutrinos, 1948
 $^{37}\text{Cl} (\nu, e^-) ^{37}\text{Ar}$

Bruno Pontecorvo
1913-1993



- Solar neutrinos, 1948
 $^{37}\text{Cl} (\nu, e^-) ^{37}\text{Ar}$
- Different ν flavors, 1957
Neutrino oscillations

Bruno Pontecorvo
1913-1993



Bruno Pontecorvo
1913-1993

- Solar neutrinos, 1948
 $^{37}\text{Cl} (\nu, e^-) ^{37}\text{Ar}$
- Different ν flavors, 1957
Neutrino oscillations
- Accelerator produced ν beams, 1959
 $\pi \rightarrow \mu + \nu, K \rightarrow \mu + \nu$

Solar Neutrinos

Solar Neutrinos

- First experiment to detect solar neutrinos by Ray Davis et al. relied on Pontecorvo's idea

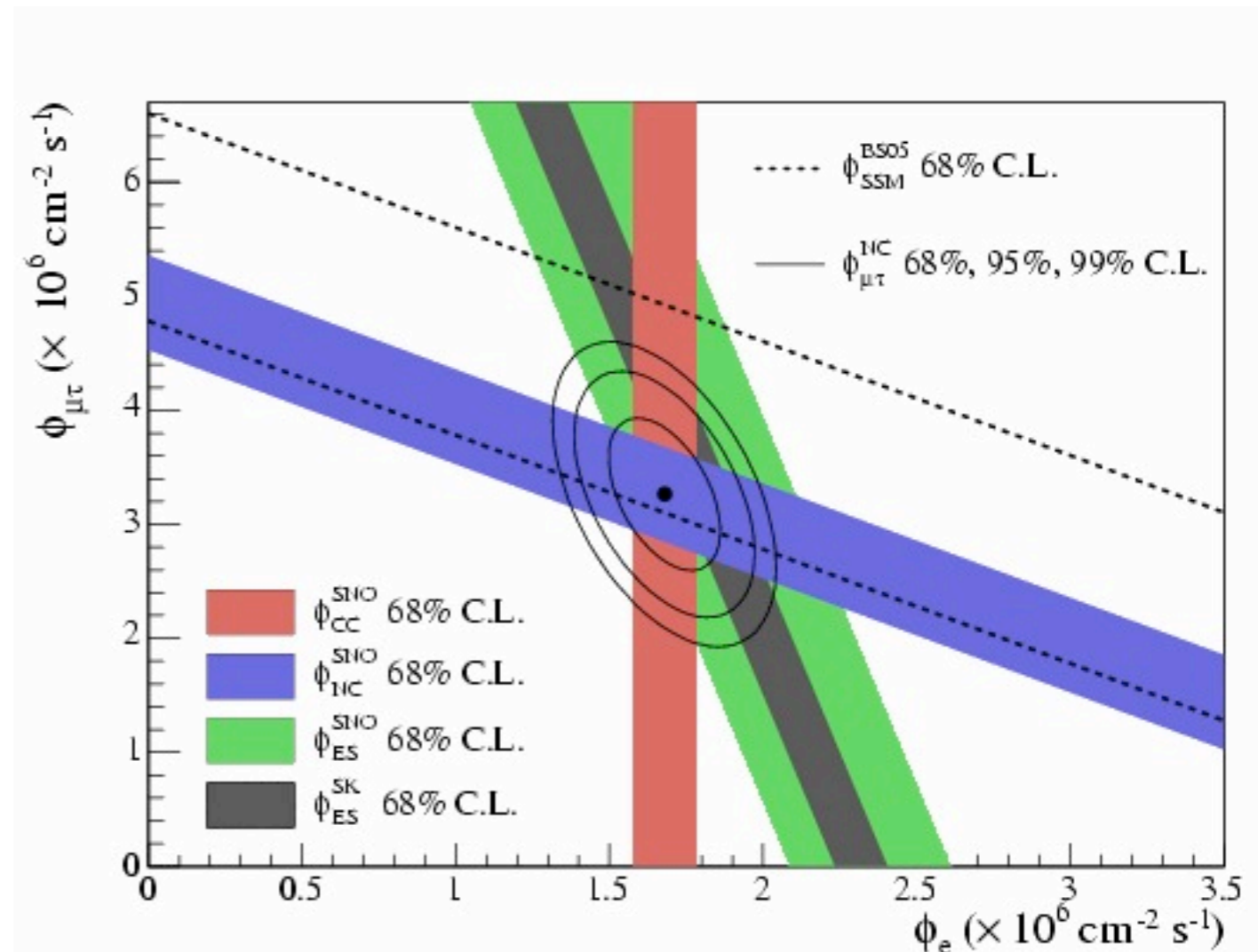
Solar Neutrinos

- First experiment to detect solar neutrinos by Ray Davis et al. relied on Pontecorvo's idea
- Nobel Prize for "Neutrino Astronomy" with M.Koshihara (2002)

Solar Neutrinos

- First experiment to detect solar neutrinos by Ray Davis et al. relied on Pontecorvo's idea
- Nobel Prize for "Neutrino Astronomy" with M.Koshihara (2002)

Resolution of the solar neutrino puzzle by SNO and other experiments



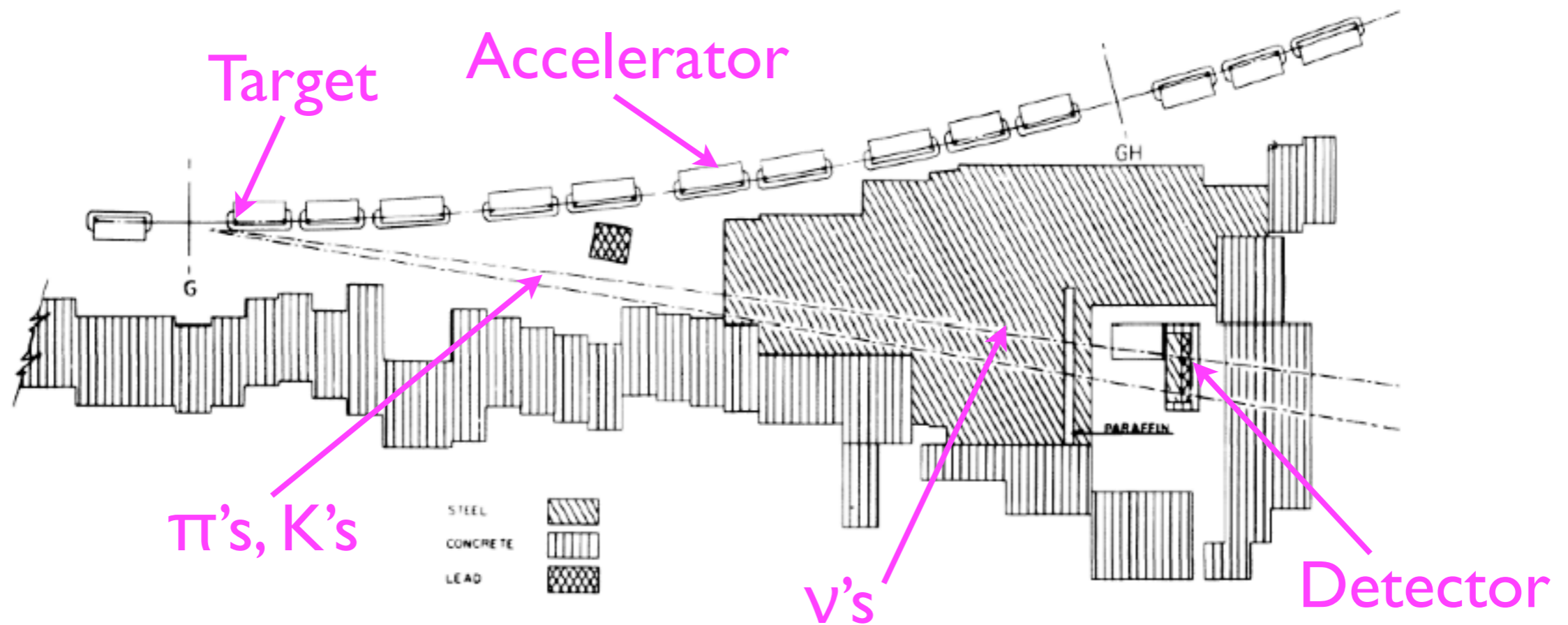
Accelerator Neutrinos

Accelerator Neutrinos

- Seminal idea (independently also by M.Schwartz) to produce high energy neutrino beams

Accelerator Neutrinos

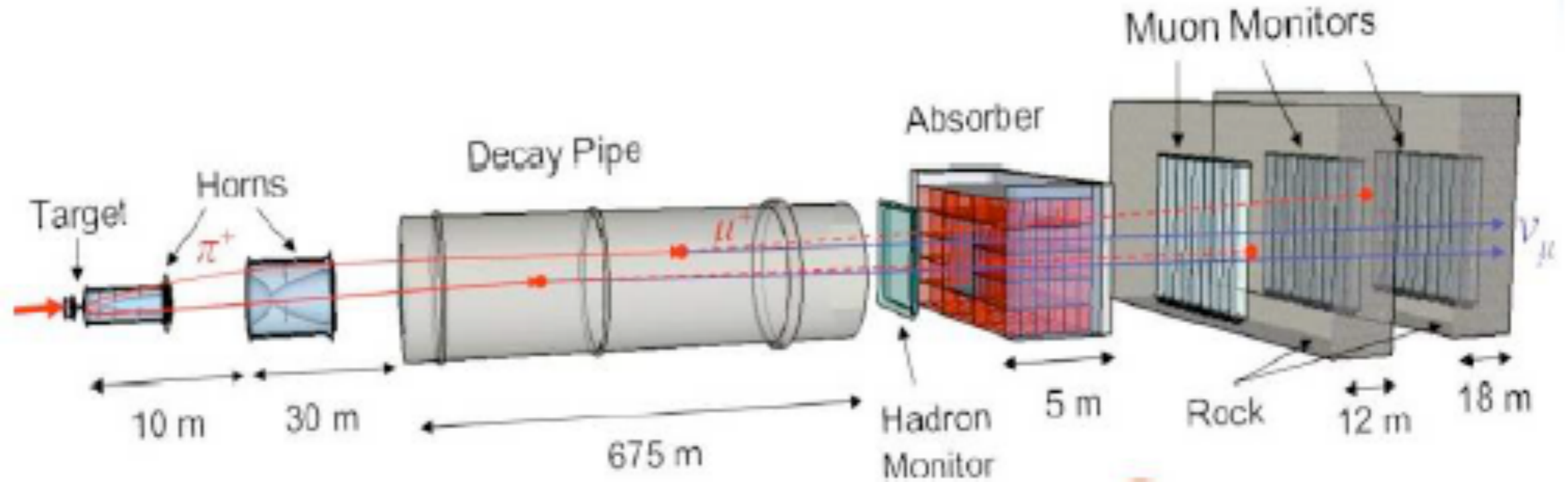
- Seminal idea (independently also by M.Schwartz) to produce high energy neutrino beams



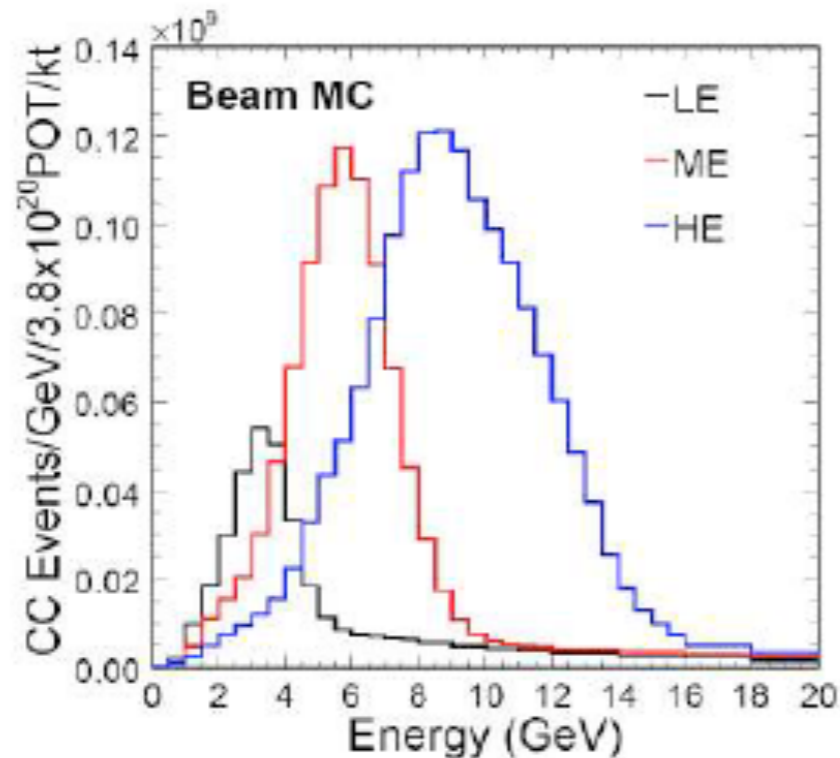
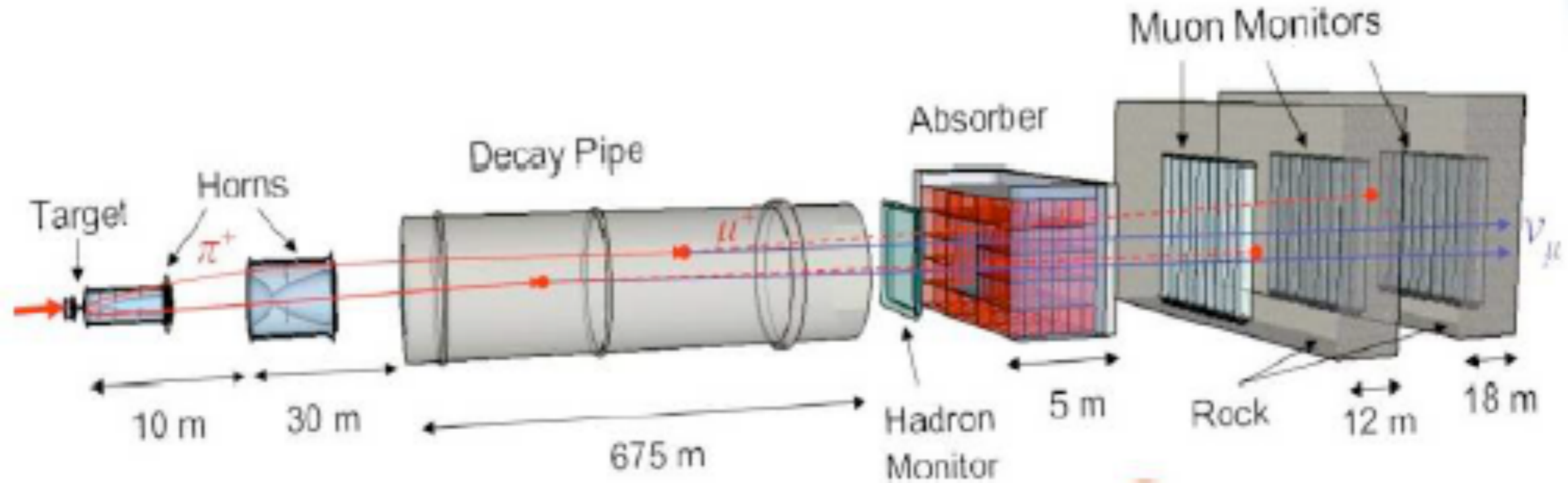
The beam for 2-neutrino experiment (BNL, 1960)

“Modern” Neutrino Beam

“Modern” Neutrino Beam

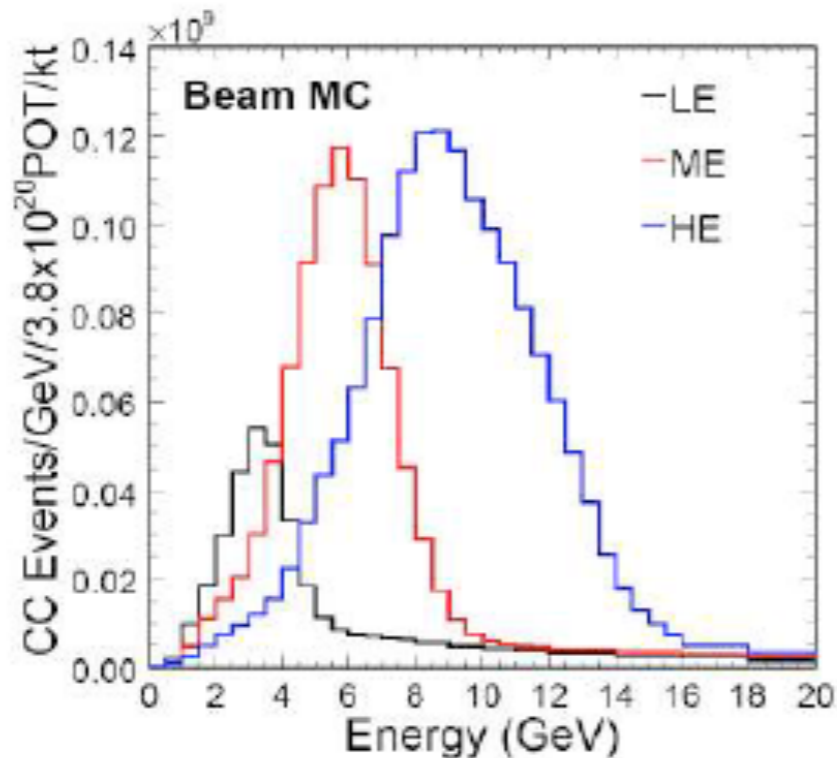
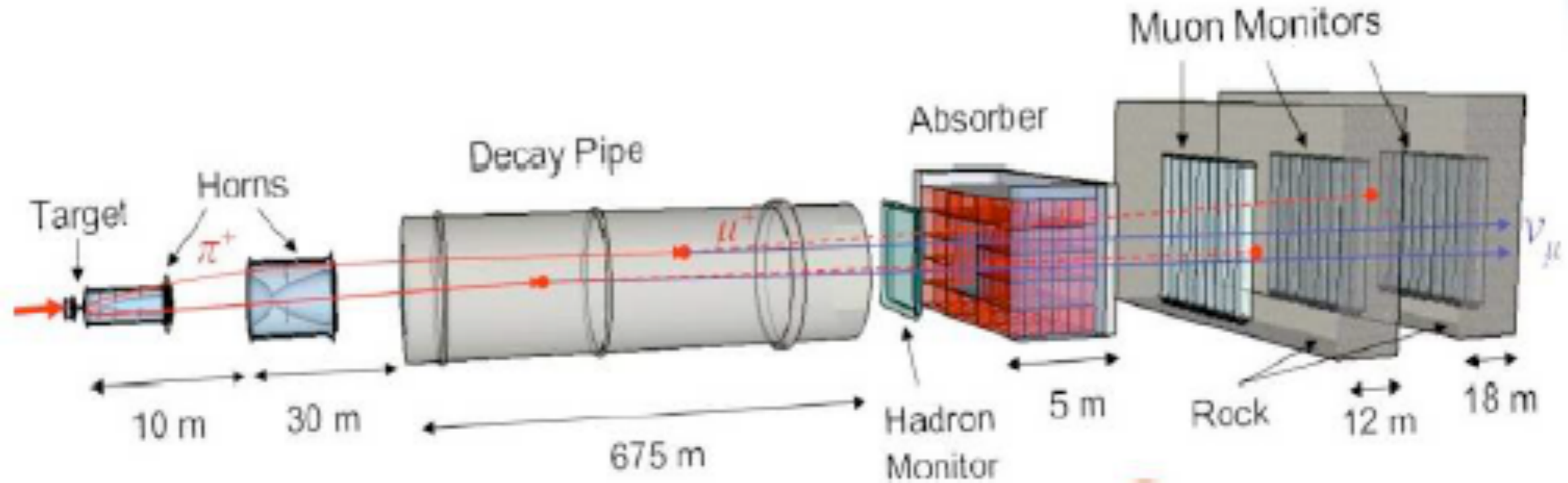


“Modern” Neutrino Beam



Ability to choose energy by moving the target

“Modern” Neutrino Beam



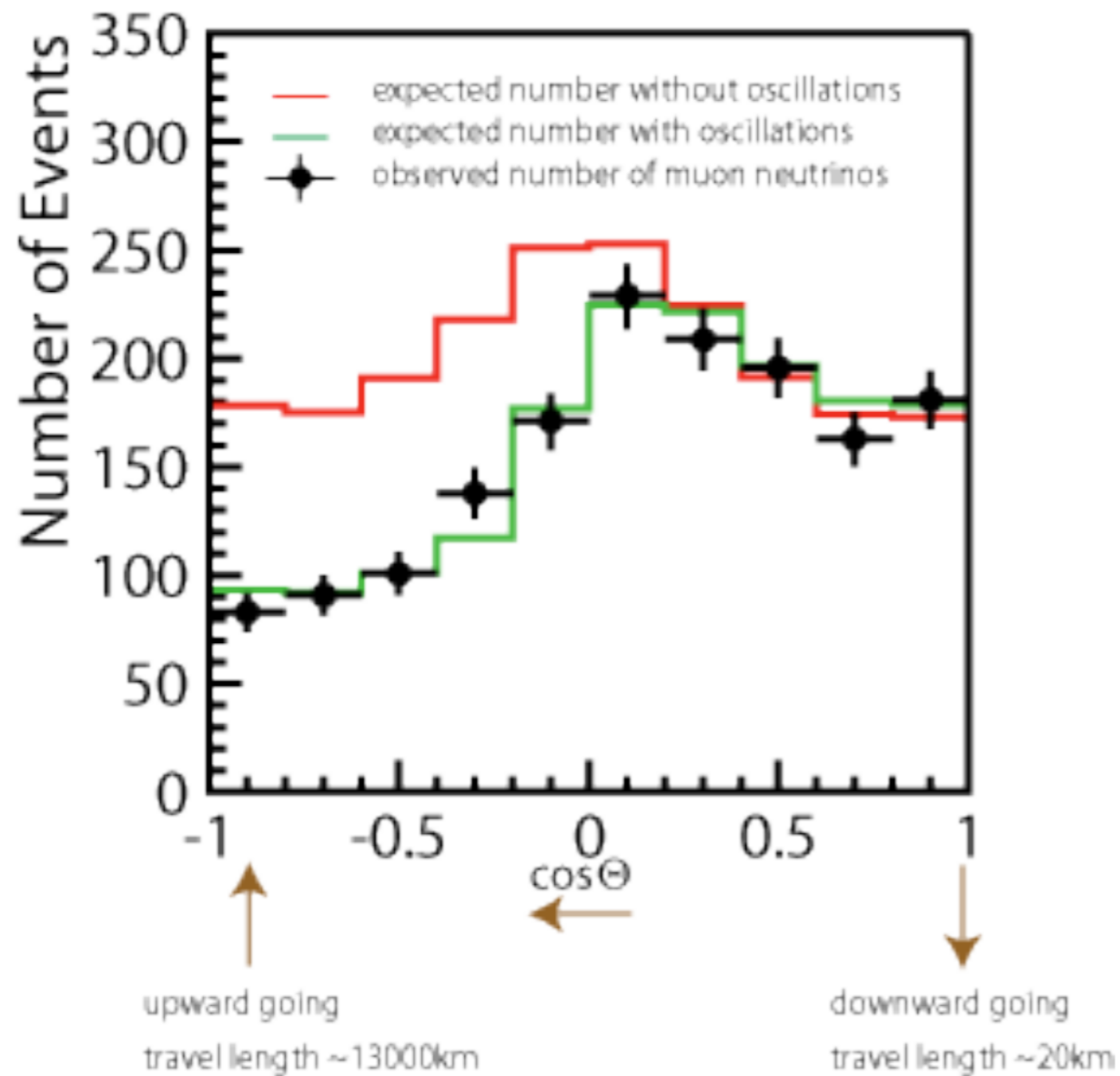
Ability to choose energy by moving the target

In the future:
Super beams \sim 2-4 Mw
Neutrino factory (μ storage ring)

Oscillations

Oscillations

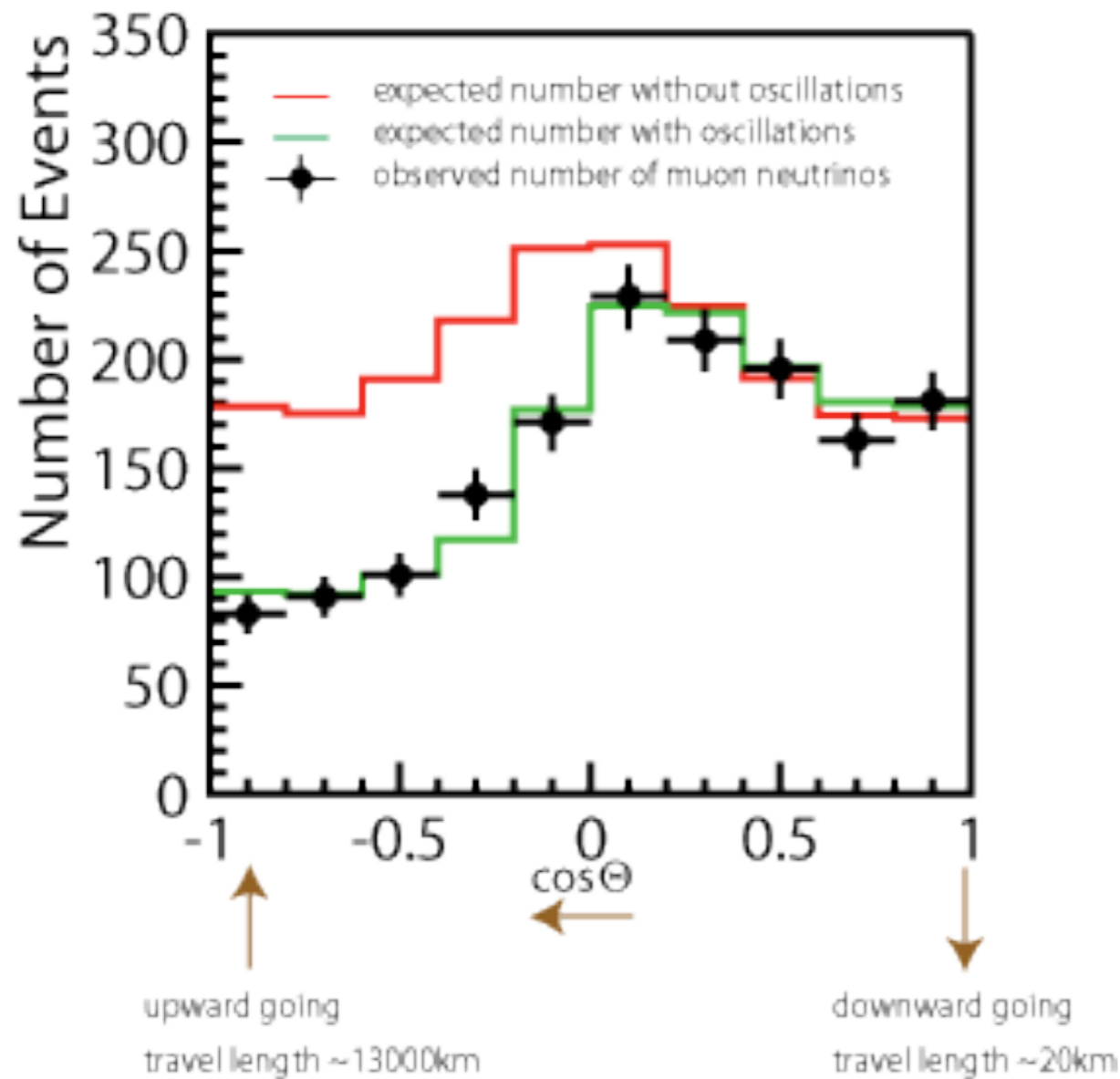
Super-Kamiokande



ν_μ 's disappear

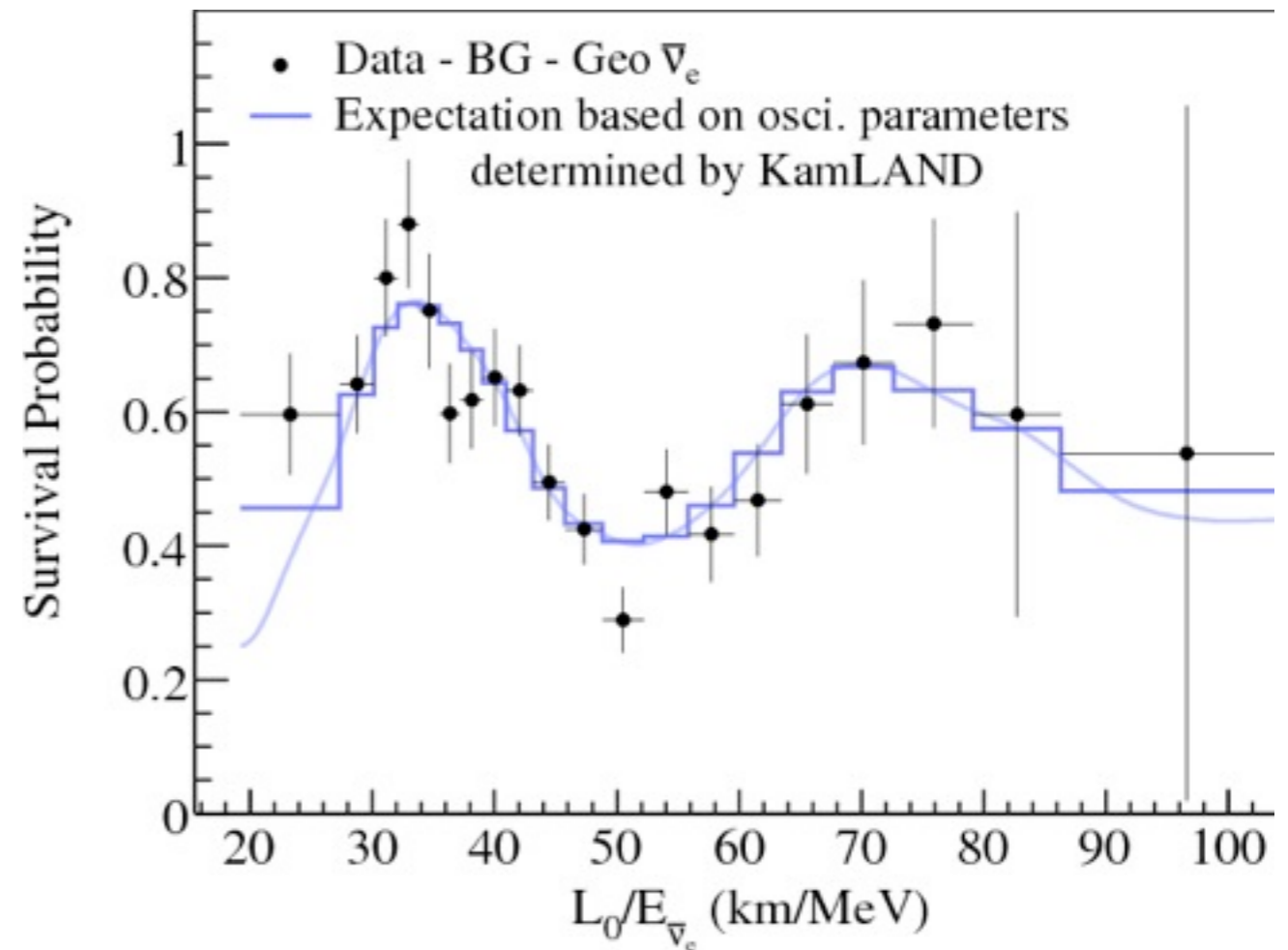
Oscillations

Super-Kamiokande



ν_μ 's disappear

KamLAND Experiment



$\bar{\nu}_e$'s disappear and then reappear

They oscillate!!

MINOS Experiment

- Oscillation Formalism
- Introduction to Experiment
(Geography, Detectors)
- Results (Oscillations)
- Prospects for the Future

Oscillations Formalism - PMNS

Oscillations Formalism - PMNS

The ν flavor and mass states are related by:

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

Oscillations Formalism - PMNS

The ν flavor and mass states are related by:

$$|\nu_\alpha\rangle = \sum_{i \in (e, \mu, \tau)} U_{\alpha i}^* |\nu_i\rangle$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Oscillations Formalism - PMNS

The ν flavor and mass states are related by:

$$|\nu_\alpha\rangle = \sum_{\alpha = (e, \mu, \tau)} \sum_i U_{\alpha i}^* |\nu_i\rangle$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31} Δm^2_{21}

Oscillations Formalism - PMNS

The ν flavor and mass states are related by:

$$|\nu_\alpha\rangle = \sum_{i \in \{e, \mu, \tau\}} U_{\alpha i}^* |\nu_i\rangle$$

$$U = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\Delta m^2_{31}} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\Delta m^2_{21}} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Oscillations are governed by 6 independent parameters:

θ_{12} , θ_{13} , θ_{23} , δ (in U matrix), and Δm^2_{21} , Δm^2_{31} ,

Oscillations Formalism - PMNS

The ν flavor and mass states are related by:

$$|\nu_\alpha\rangle = \sum_{i} U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31} Δm^2_{21}

Oscillations are governed by 6 independent parameters:

$\theta_{12}, \theta_{13}, \theta_{23}, \delta$ (in U matrix), and $\Delta m^2_{21}, \Delta m^2_{31}$,

2-flavor approximation is often adequate:

$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - \sin^2(2\theta) \sin^2(1.27 \Delta m^2 L/E)$$

What can MINOS do?

What can MINOS do?

$$|\nu_\alpha\rangle = \sum_{i} U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31}

Δm^2_{21}

What can MINOS do?

$$|\nu_\alpha\rangle = \sum_{i=1,2,3} U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31}
 Δm^2_{21}

Disappearance experiment: $\nu_\mu \rightarrow \nu_x$

What can MINOS do?

$$|\nu_\alpha\rangle = \sum_{i} U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31} Δm^2_{21}

Disappearance experiment: $\nu_\mu \rightarrow \nu_x$

Appearance experiment: $\nu_\mu \rightarrow \nu_e$

What can MINOS do?

$$|\nu_\alpha\rangle = \sum_{i} U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31}
 Δm^2_{21}

Disappearance experiment: $\nu_\mu \rightarrow \nu_x$

Appearance experiment: $\nu_\mu \rightarrow \nu_e$

CPT, Anomalous interactions: $\bar{\nu}_\mu \rightarrow \nu_x$

What can MINOS do?

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

$\alpha = (e, \mu, \tau, s)$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Δm^2_{31} Δm^2_{21}

Disappearance experiment: $\nu_\mu \rightarrow \nu_x$

Appearance experiment: $\nu_\mu \rightarrow \nu_e$

CPT, Anomalous interactions: $\bar{\nu}_\mu \rightarrow \nu_x$

Search for a 4th, sterile neutrino: $\nu_\mu \rightarrow \nu_s$

MINOS Geography

MINOS Geography



MINOS Geography



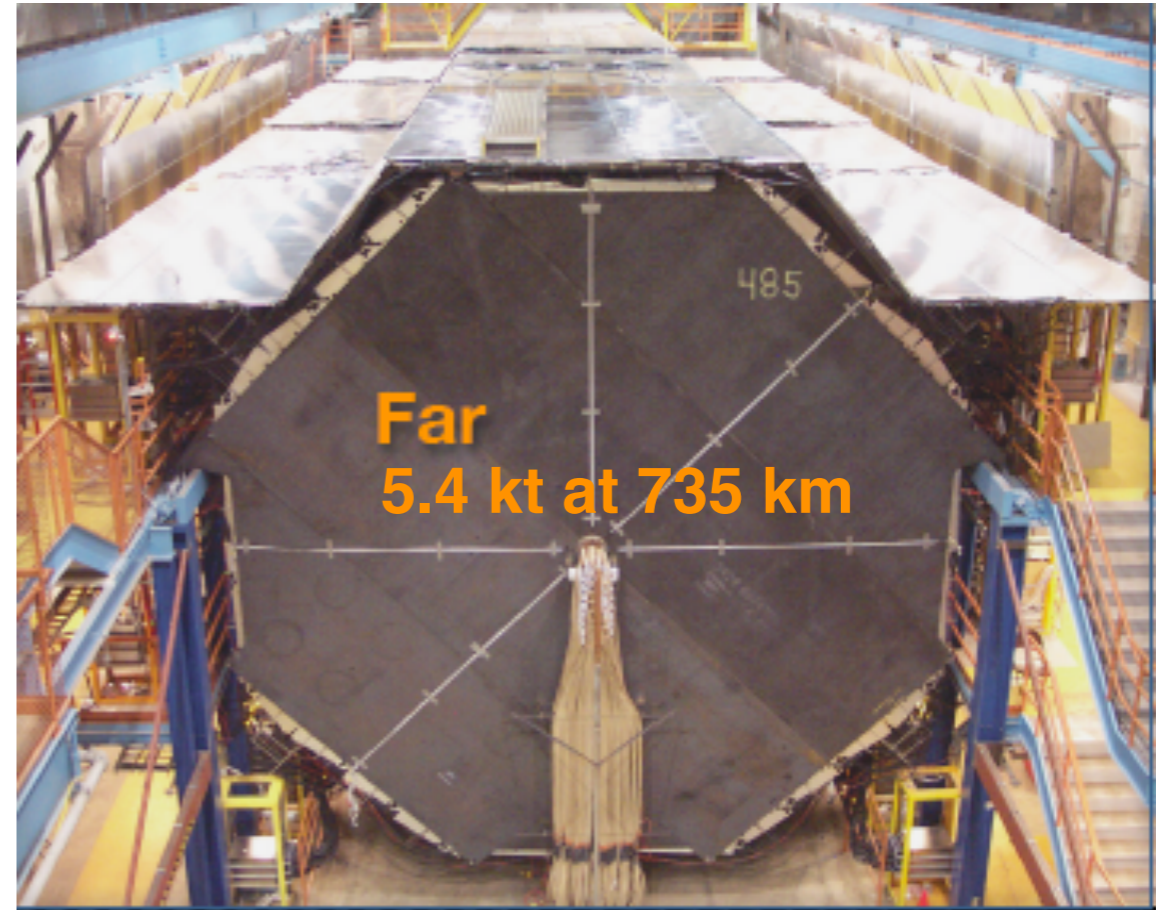
The MINOS Collaboration



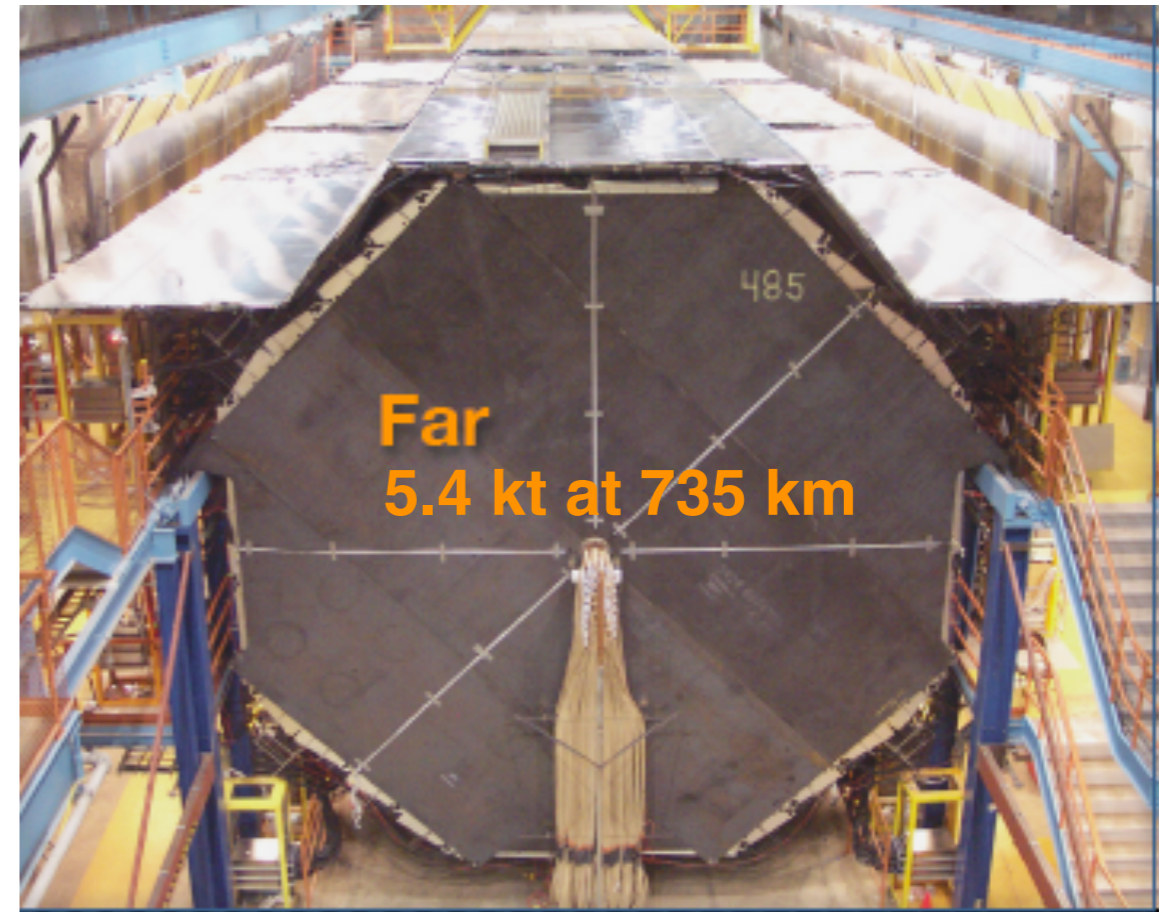
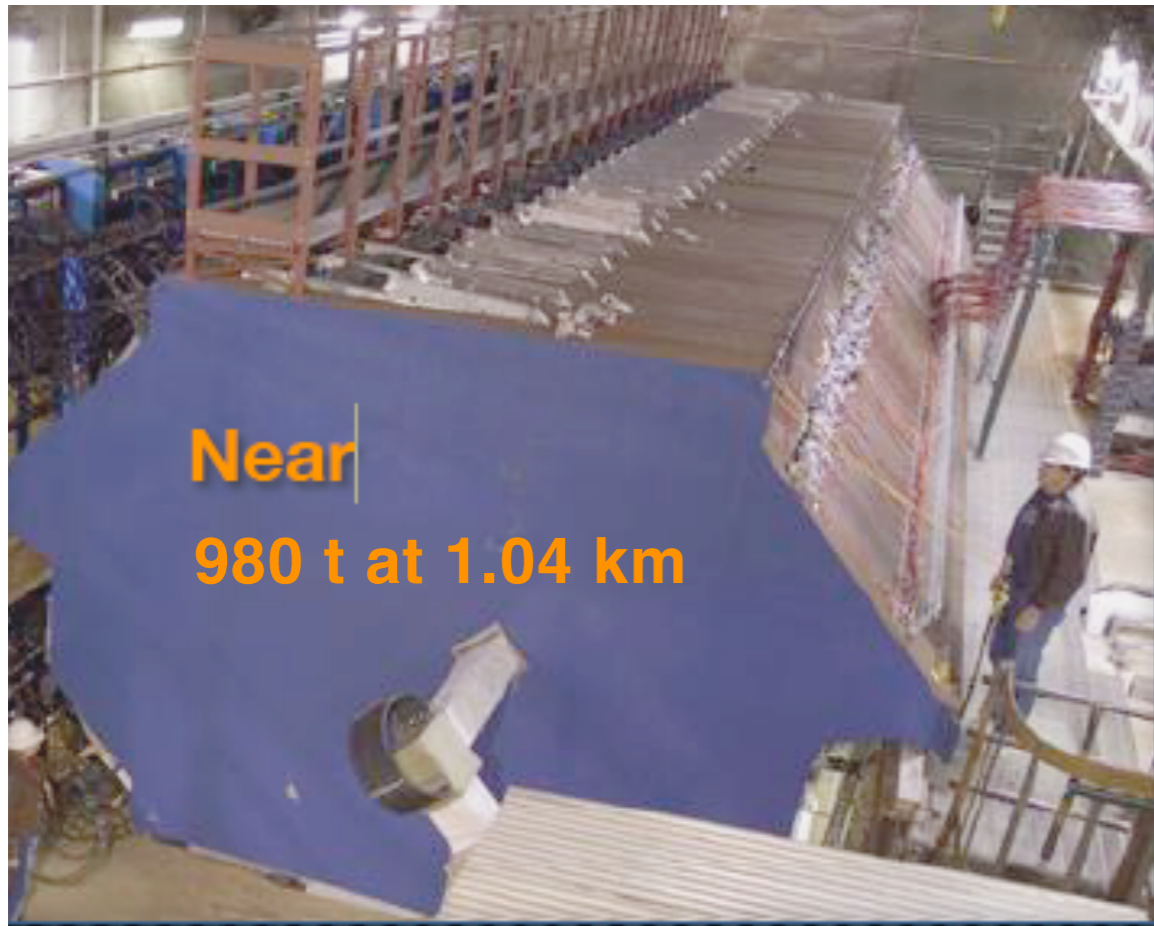
Argonne • Athens • Benedictine • Brookhaven • Caltech • Cambridge • Campinas • Fermilab
Goias • Harvard • Holy Cross • IIT • Indiana • Iowa State • Minnesota-Twin Cities
Minnesota-Duluth • Otterbein • Oxford • Pittsburgh • Rutherford • Sao Paulo • South Carolina • Stanford • Sussex • Texas
A&M • Texas-Austin • Tufts • UCL • Warsaw • William & Mary

MINOS Detectors

MINOS Detectors

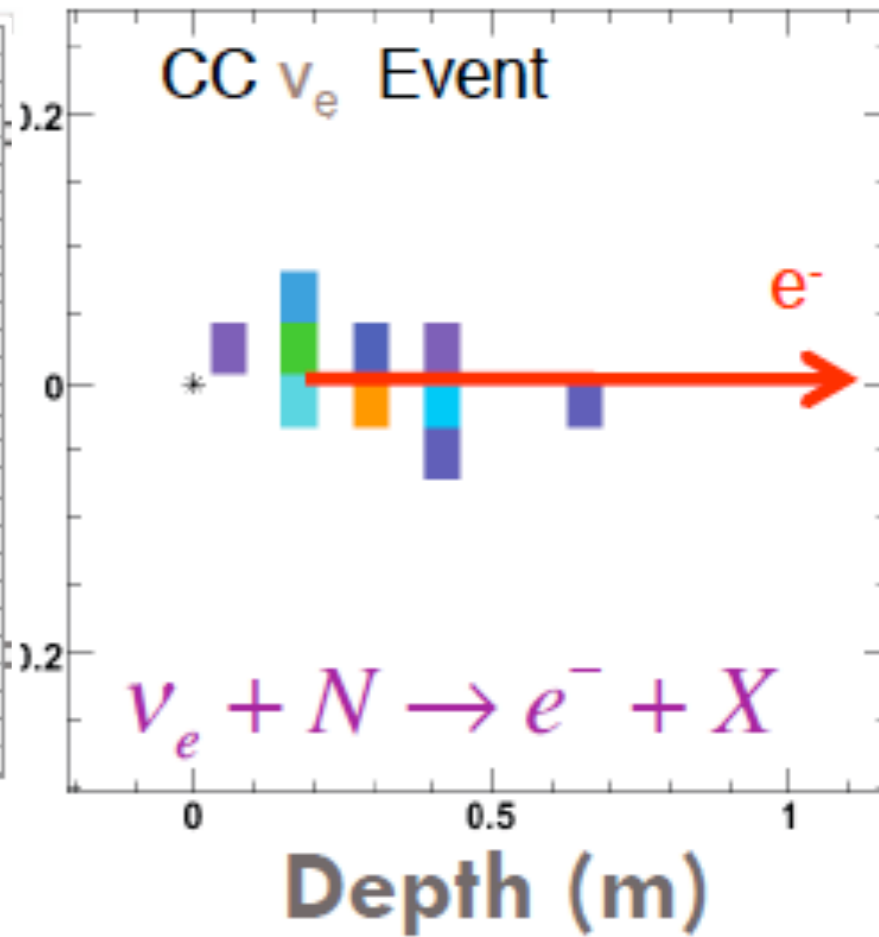
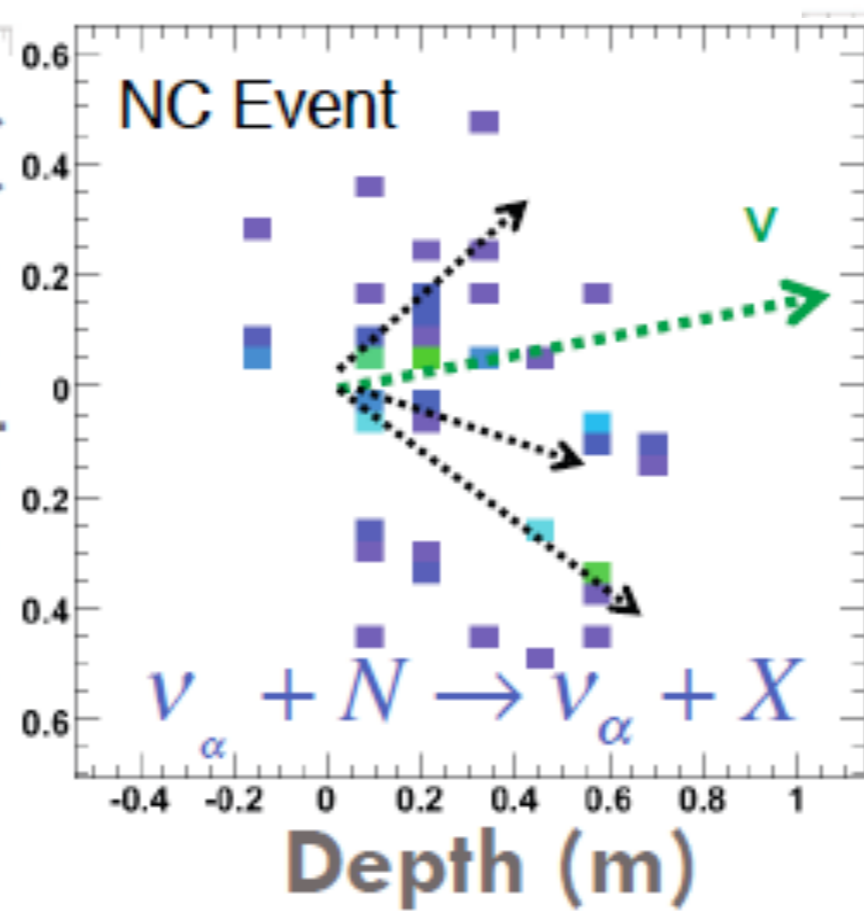
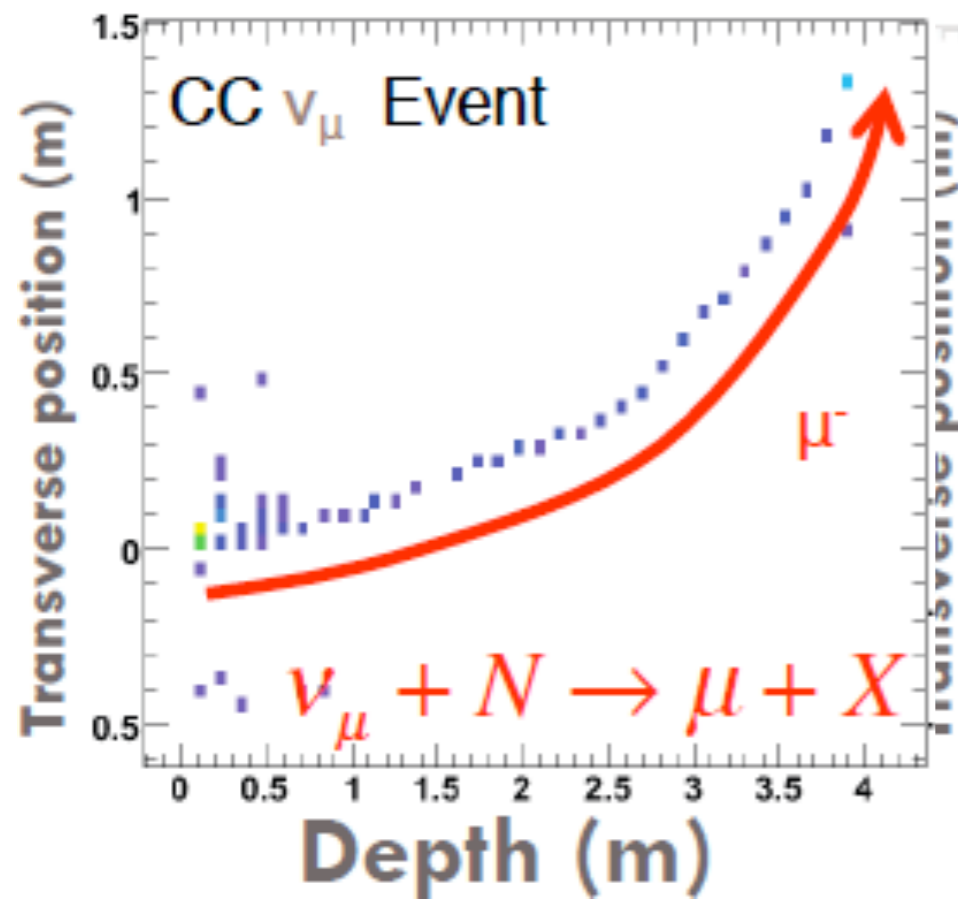


MINOS Detectors



- As similar as possible functionally
- Alternating layers of steel (2.5 cm thick) and scintillator
- Alternating scintillator planes at 90 deg, 4.1 cm strips
- Light collection by wavelength shifting fibers
- Readout by 64 ch(ND) or 16 ch(FD) multi-anode PMT's
- Magnetized, average B field 1.3 T

Types of Events



Disappearance (ν_μ ; $\theta_{23}, \Delta m^2_{31}$)

Disappearance ($\nu_\mu; \theta_{23}, \Delta m^2_{31}$)

Relatively easy: long track

Good energy measurement:

Calorimetry for hadronic shower

Range or curvature for muon

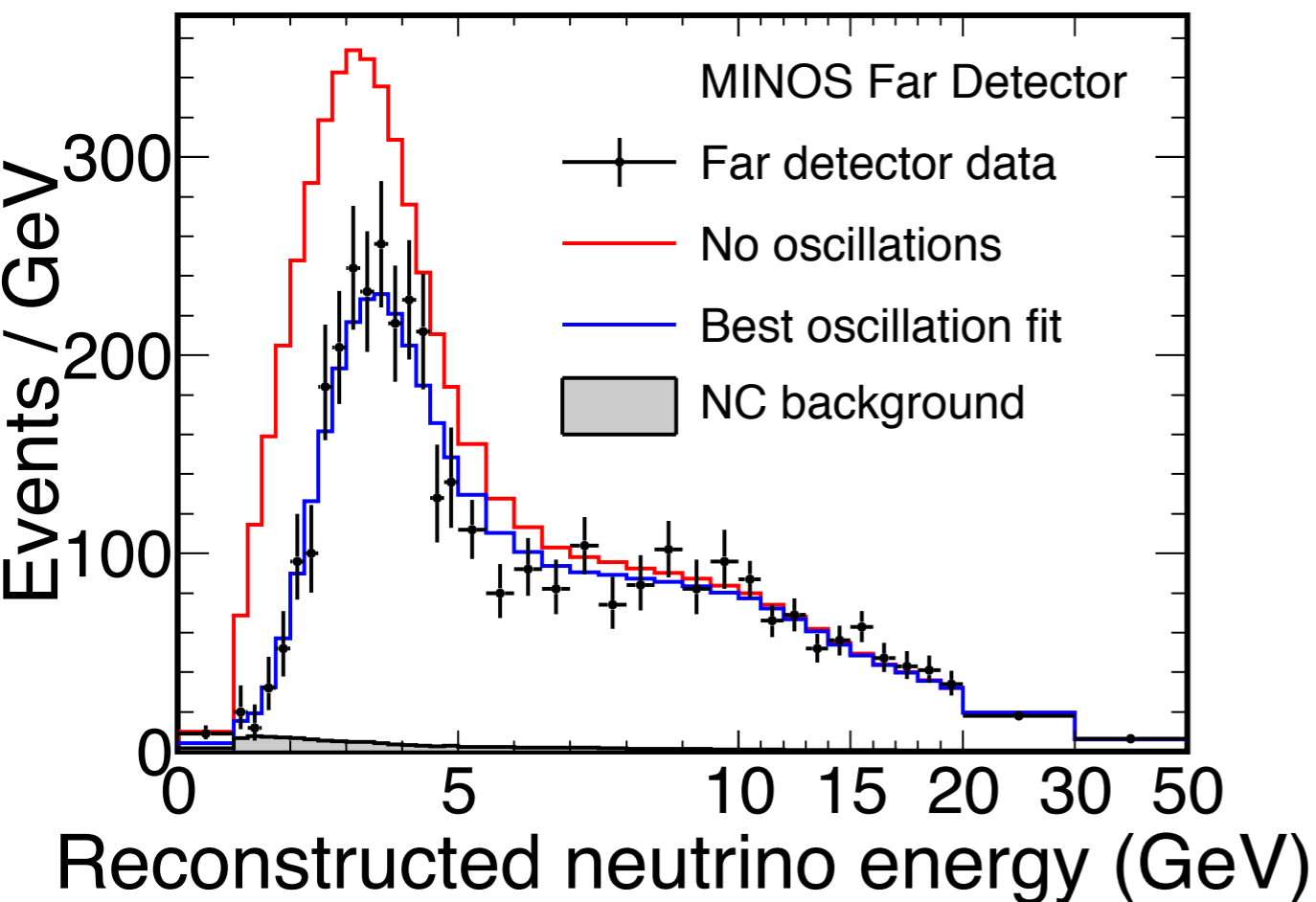
Disappearance ($\nu_\mu; \theta_{23}, \Delta m^2_{31}$)

Relatively easy: long track

Good energy measurement:

Calorimetry for hadronic shower

Range or curvature for muon



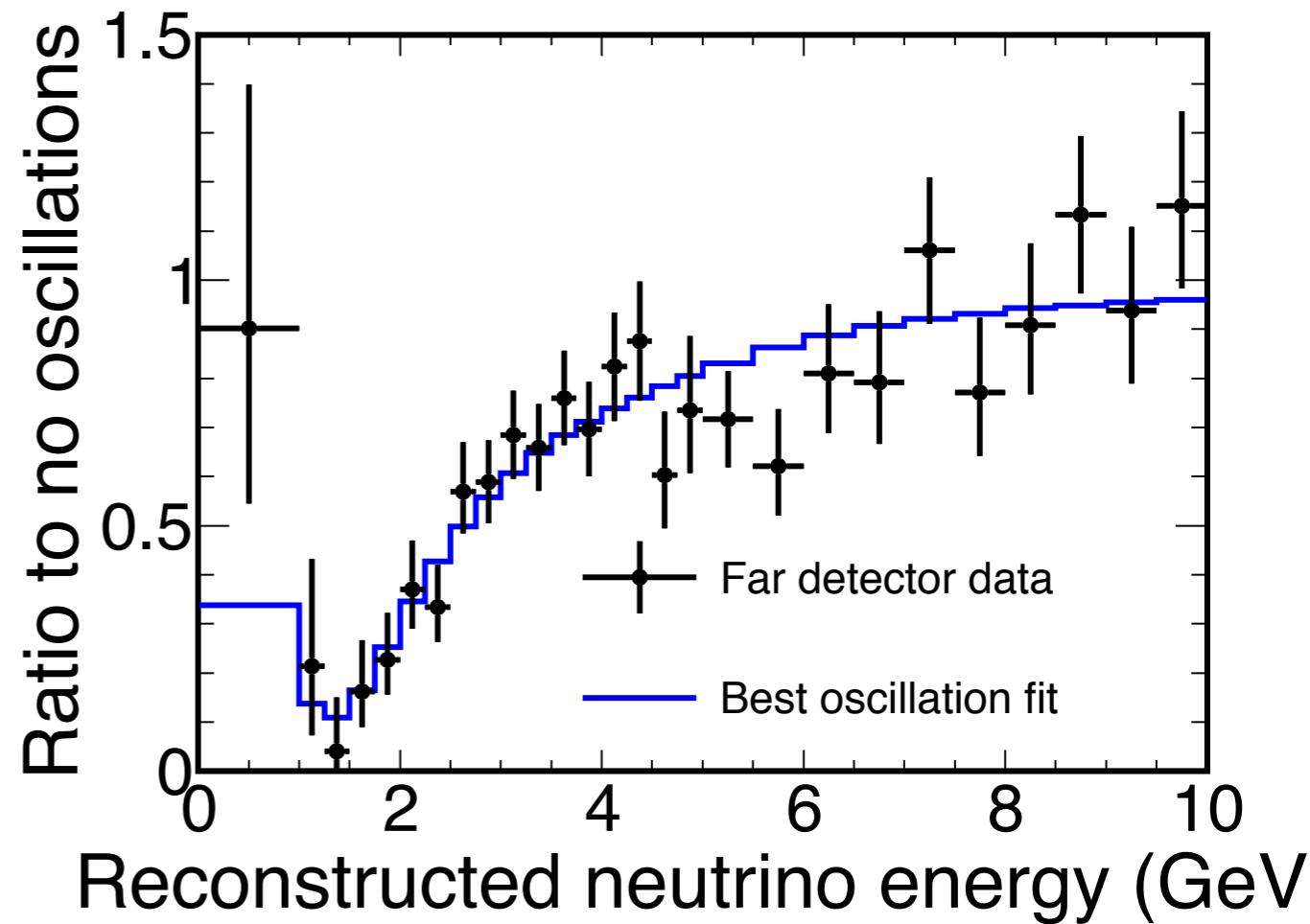
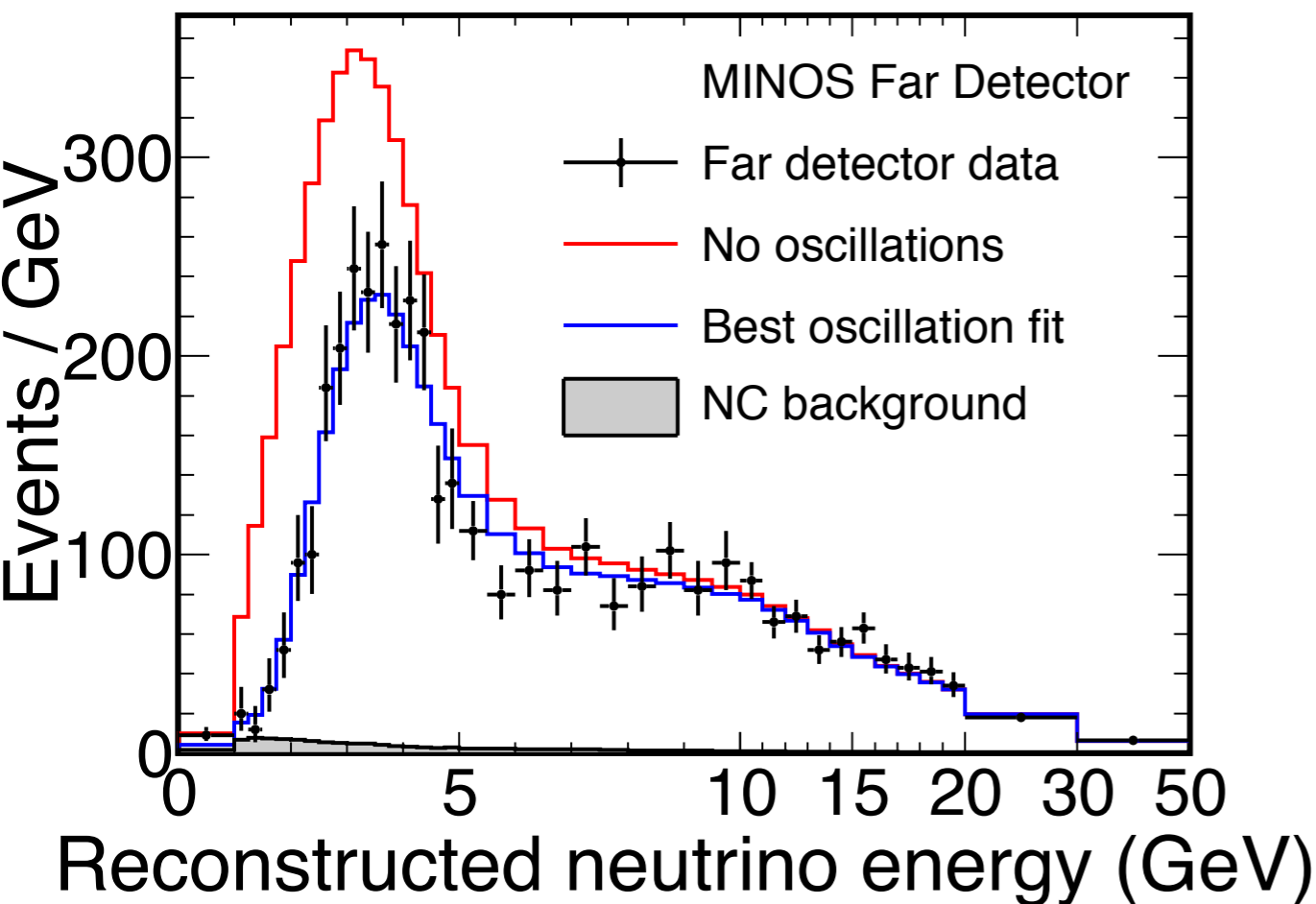
Disappearance ($\nu_\mu; \theta_{23}, \Delta m^2_{31}$)

Relatively easy: long track

Good energy measurement:

Calorimetry for hadronic shower

Range or curvature for muon



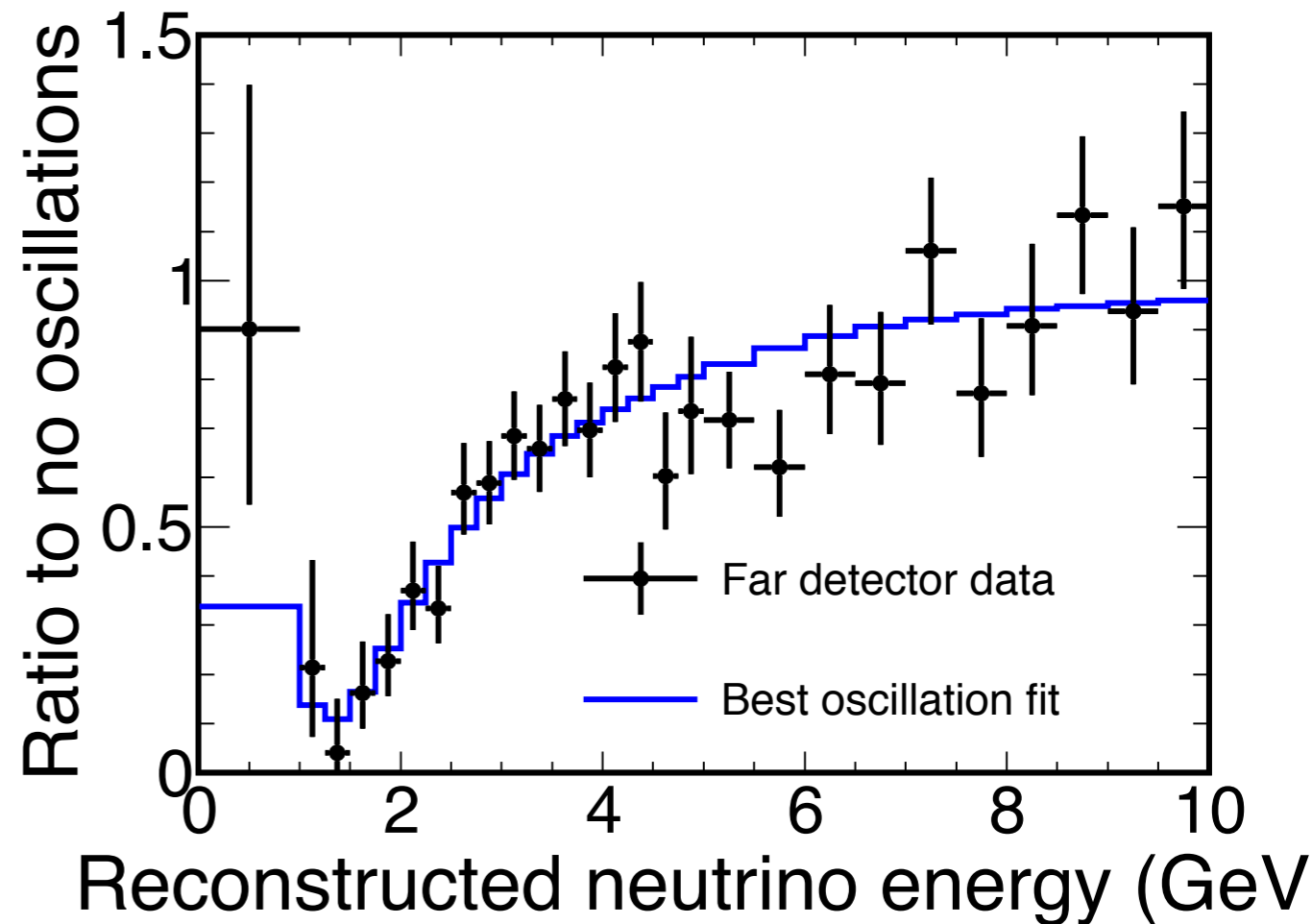
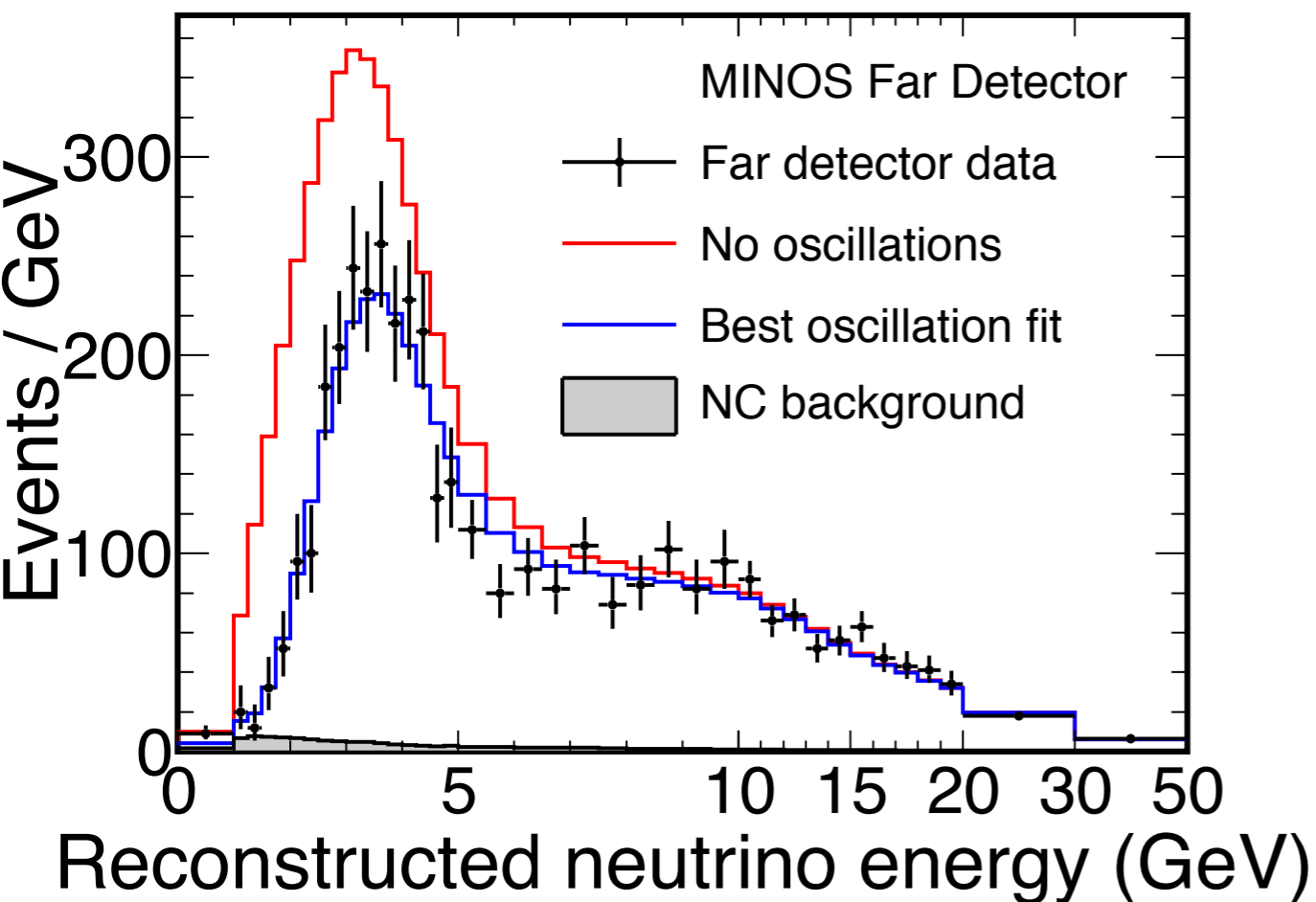
Disappearance ($\nu_\mu; \theta_{23}, \Delta m^2_{31}$)

Relatively easy: long track

Good energy measurement:

Calorimetry for hadronic shower

Range or curvature for muon



Result

$$\Delta m^2 = 2.32^{+0.12}_{-0.08} \times 10^{-3} \text{ eV}^2$$

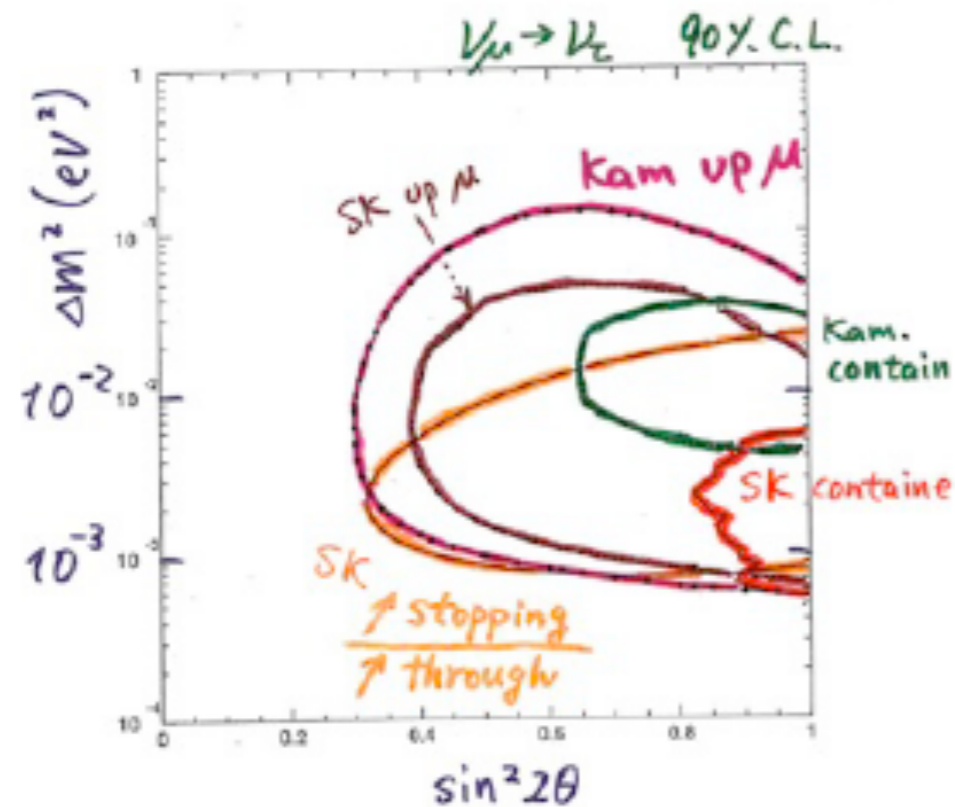
$$\sin^2(2\theta) > 0.90$$

Disappearance: 1998 and 2012

Disappearance: 1998 and 2012

Summary

Evidence for ν_μ oscillations



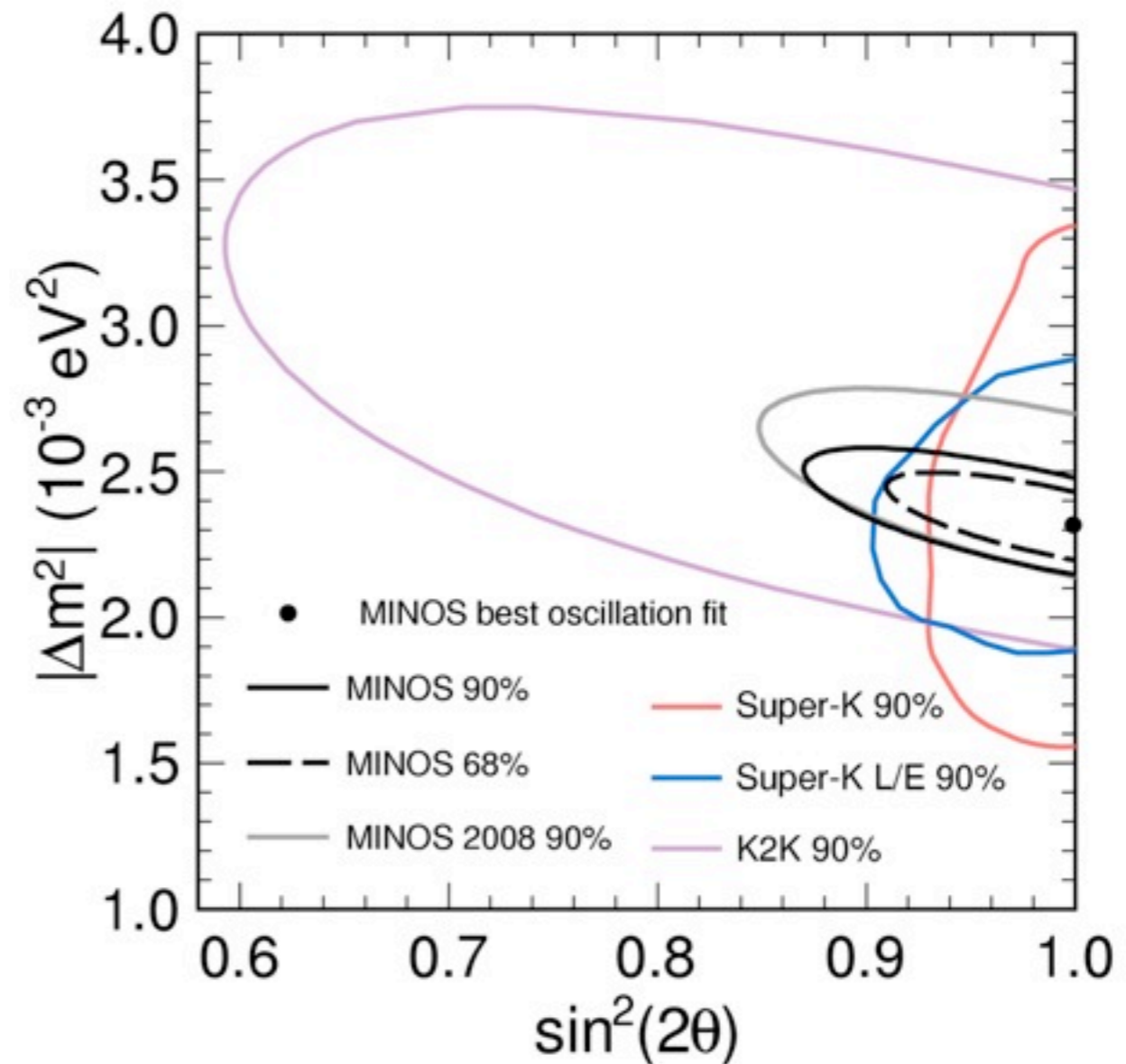
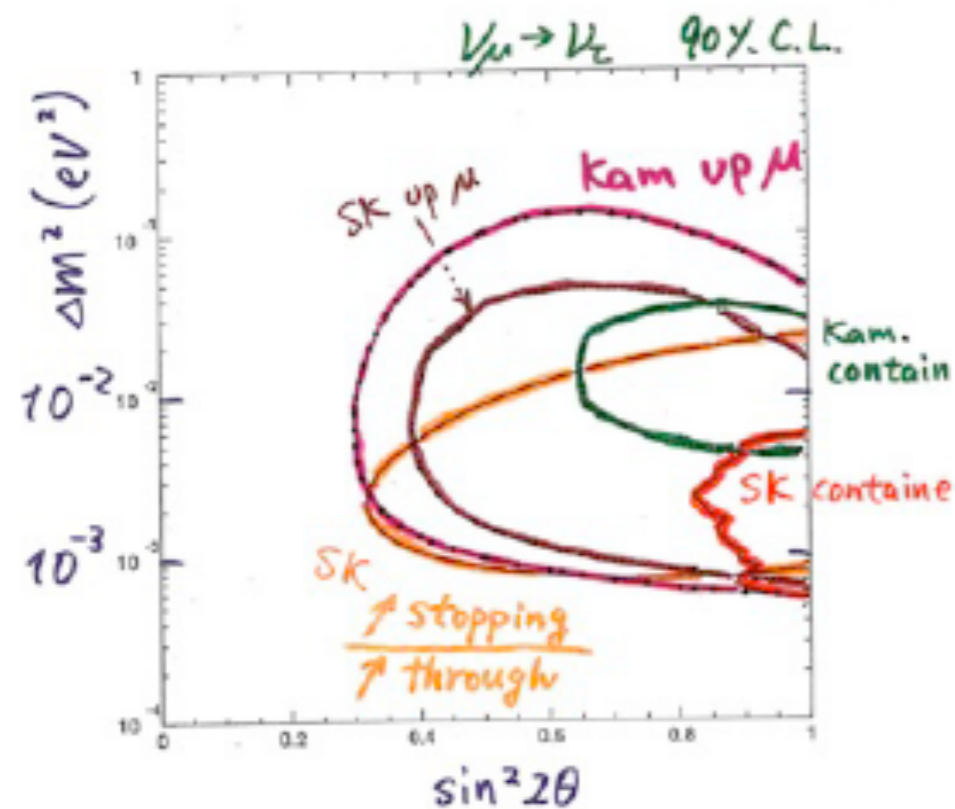
- $\begin{cases} \sin^2 2\theta > 0.8 \\ \Delta m^2 \sim 10^{-3} \sim 10^{-2} \end{cases}$

(• $\nu_\mu \rightarrow \nu_e$ or $\nu_\mu \rightarrow \nu_s$?)

Disappearance: 1998 and 2012

Summary

Evidence for ν_μ oscillations



- $\begin{cases} \sin^2 2\theta > 0.8 \\ \Delta m^2 \sim 10^{-3} \sim 10^{-2} \end{cases}$

(• $\nu_\mu \rightarrow \nu_e$ or $\nu_\mu \rightarrow \nu_s$?)

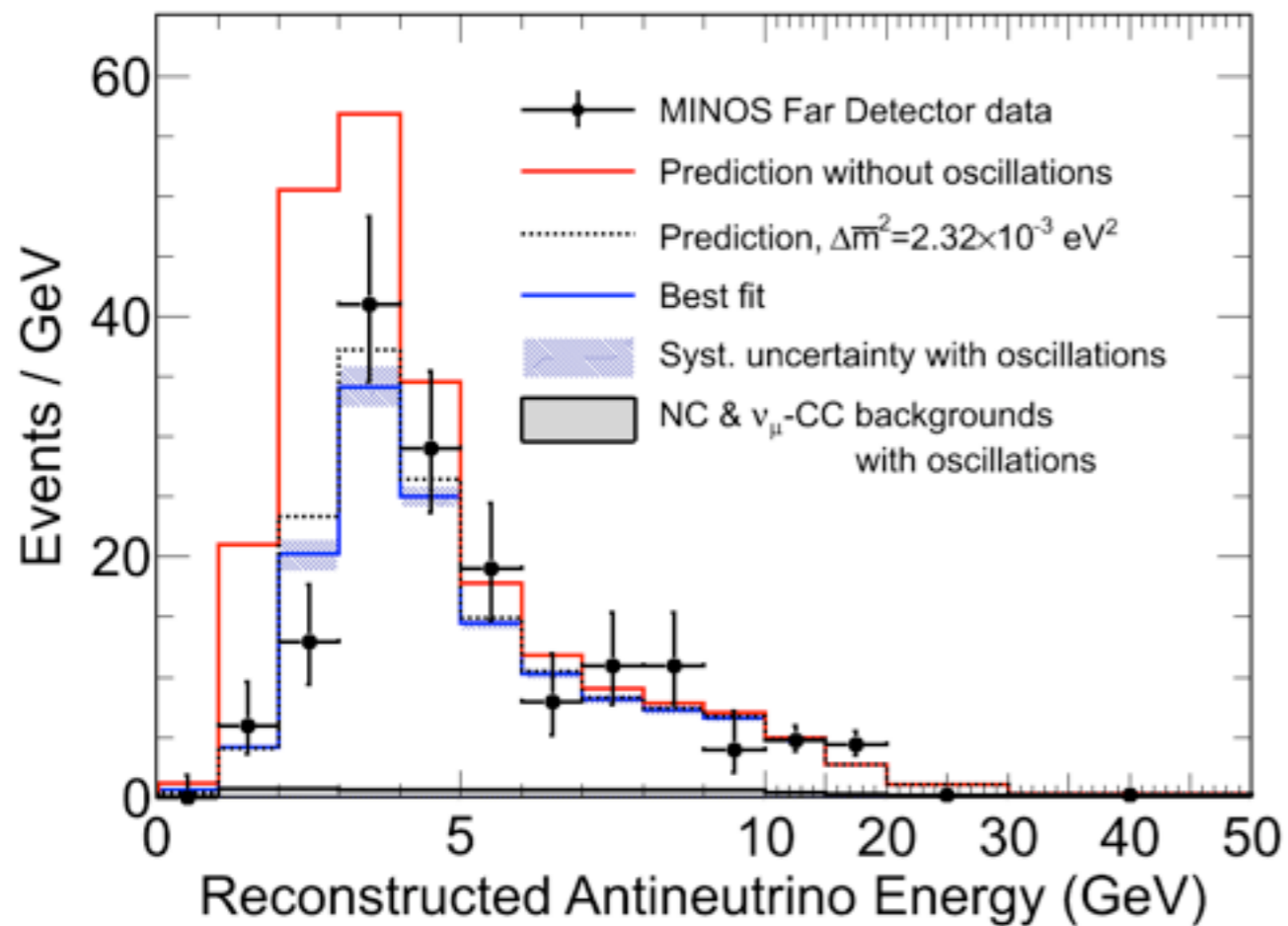
Result

$$\Delta m^2 = 2.32^{+0.12}_{-0.08} \times 10^{-3} \text{ eV}^2$$

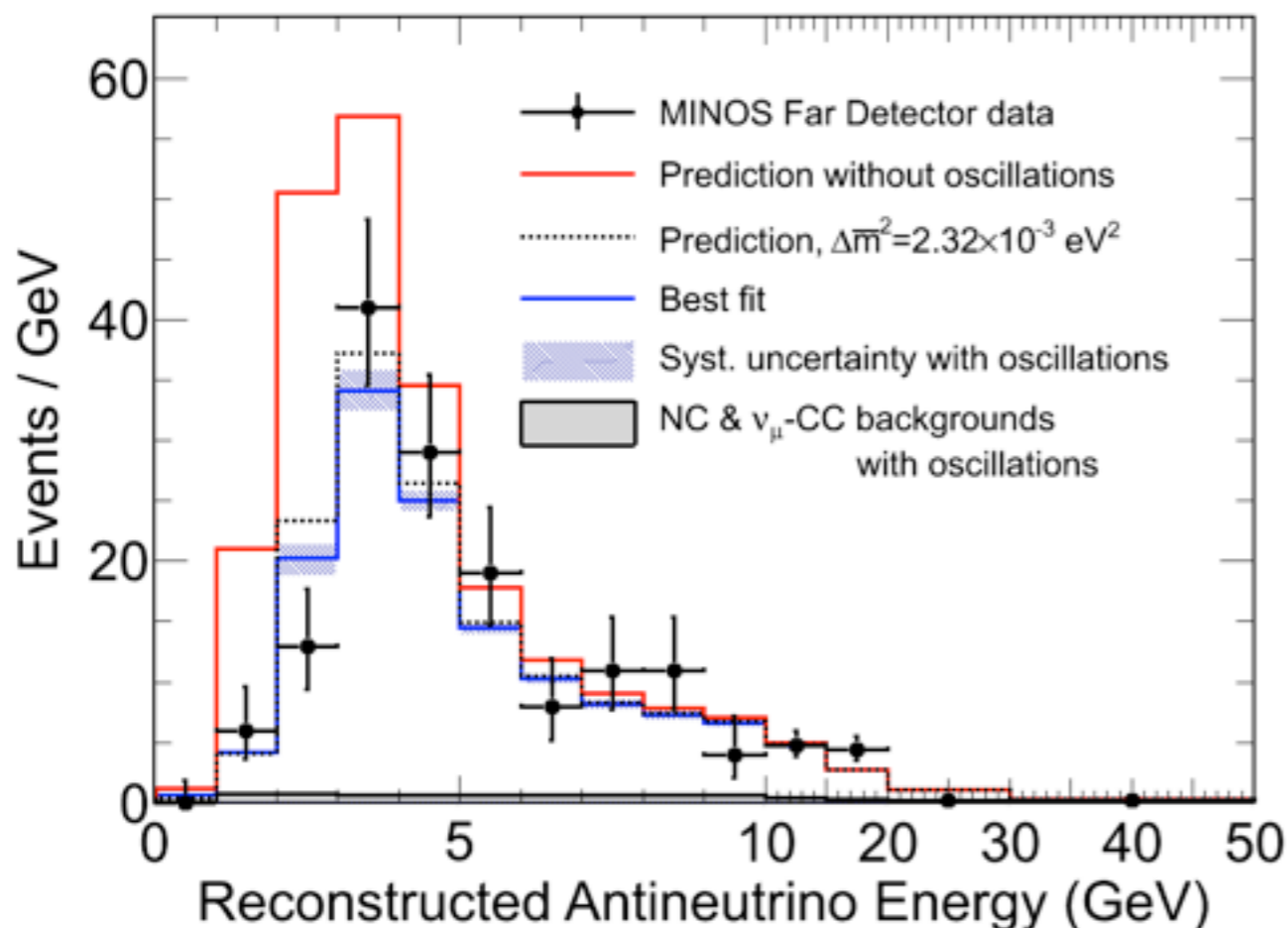
$$\sin^2(2\theta) > 0.90$$

Disappearance (\overline{V}_μ)

Disappearance ($\bar{\nu}_\mu$)



Disappearance ($\bar{\nu}_\mu$)

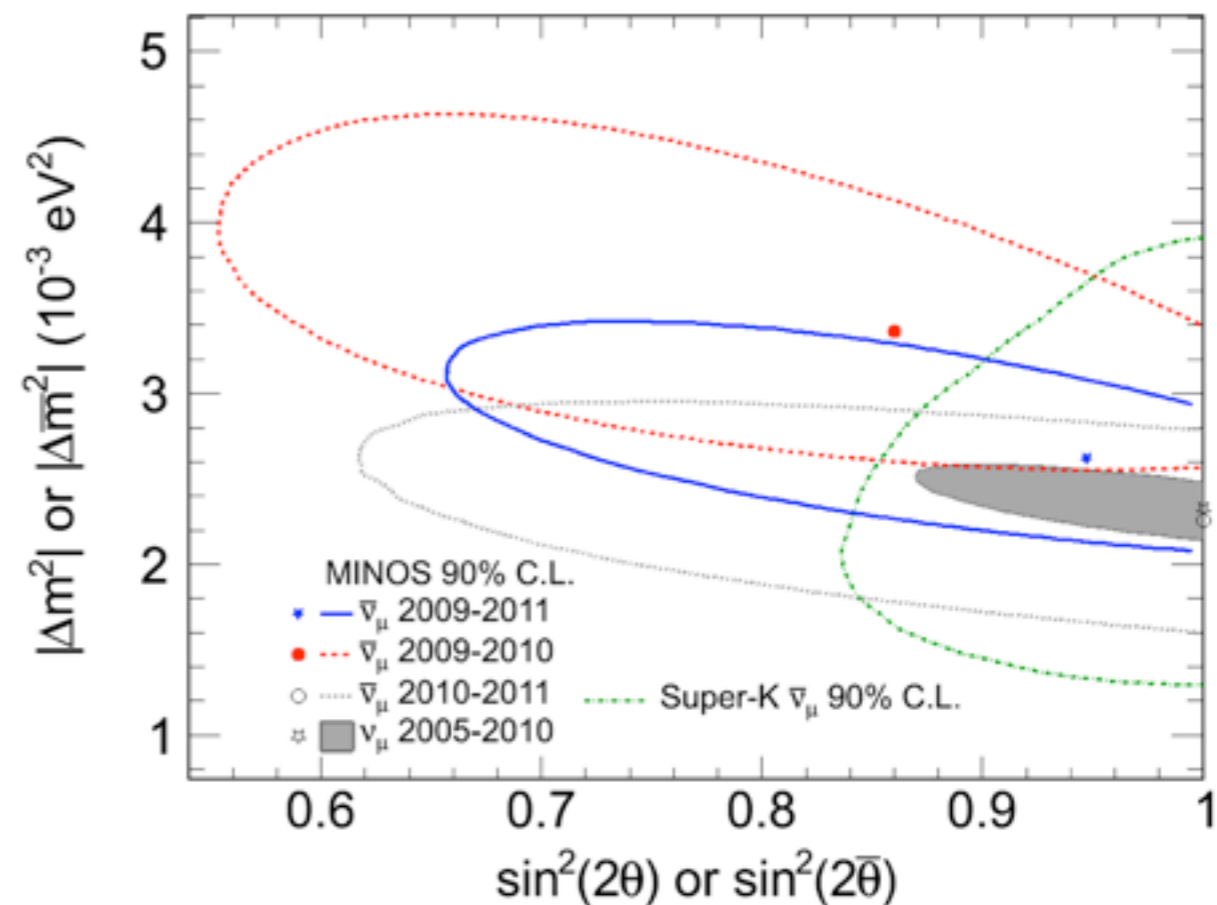
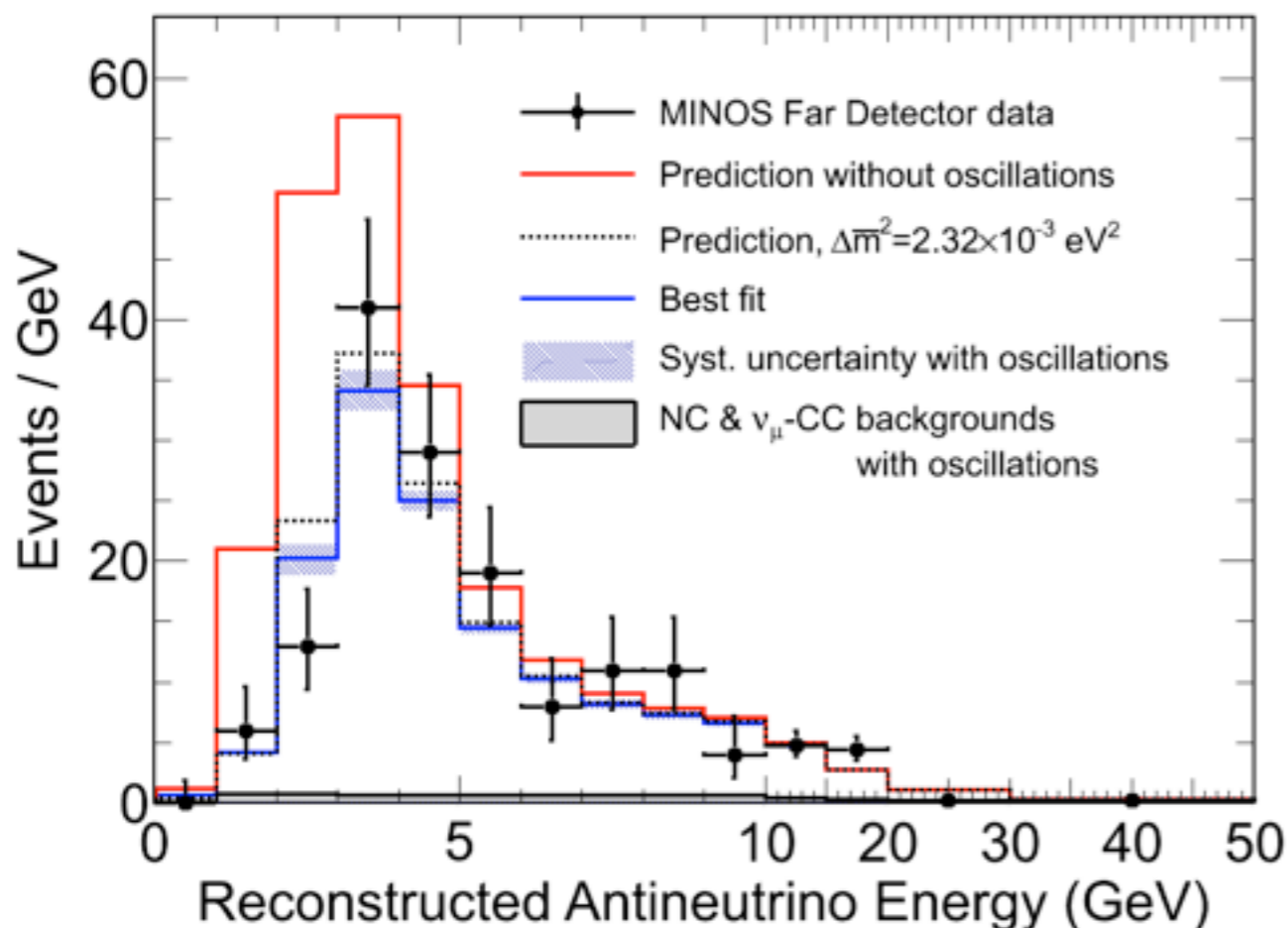


Result

$$\Delta\bar{m}^2=2.62^{+0.31}_{-0.28}(\text{stat.})\pm 0.09(\text{sys}) \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.750$$

Disappearance ($\bar{\nu}_\mu$)



Result

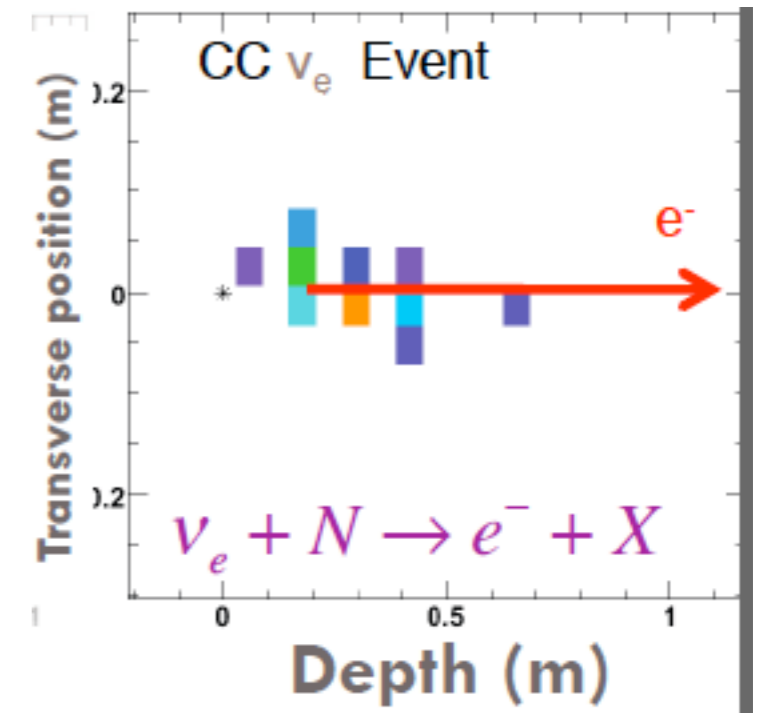
$$\Delta\bar{m}^2 = 2.62^{+0.31}_{-0.28}(\text{stat.}) \pm 0.09(\text{sys}) \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.750$$

Appearance ($v_e; \theta_{13}$)

Appearance ($\nu_e; \theta_{13}$)

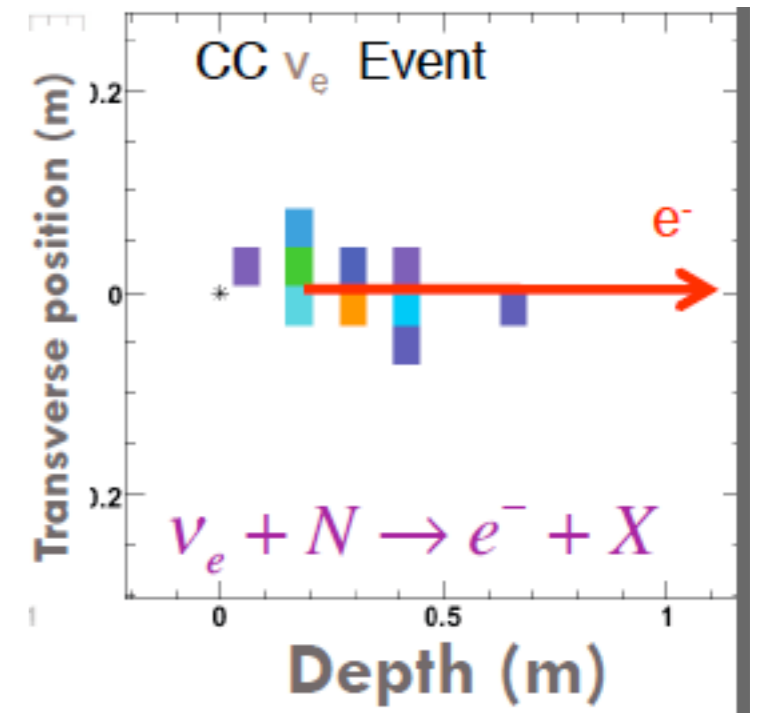
Events characterized by a relatively short and narrow shower



Appearance (ν_e ; θ_{13})

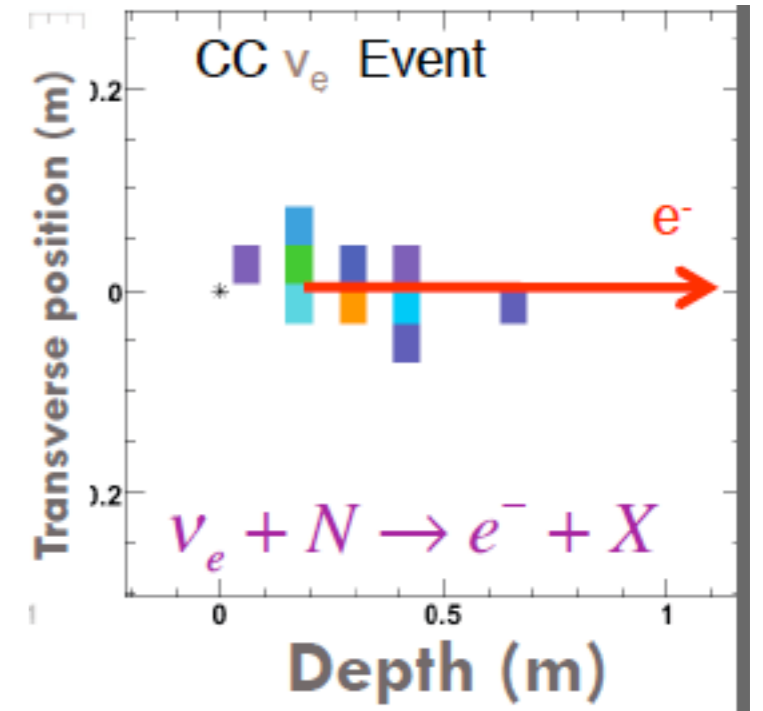
Events characterized by a relatively short and narrow shower

Separation from more numerous NC background done on statistical basis



Appearance ($\nu_e; \theta_{13}$)

Events characterized by a relatively short and narrow shower
 Separation from more numerous NC background done on statistical basis



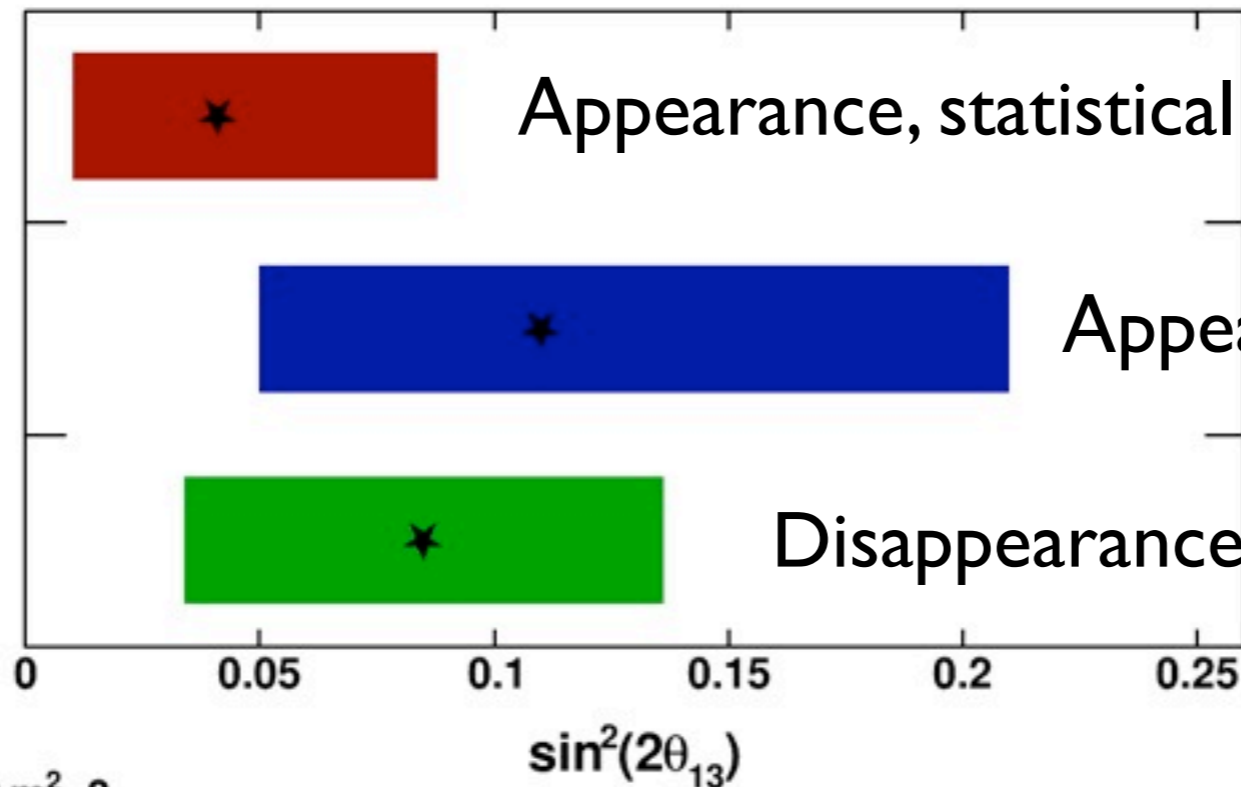
Current Status

68% CL Allowed

MINOS
(PRL107.181802)

T2K
(PRL107.041801)

Double Chooz
(LowNu2011)

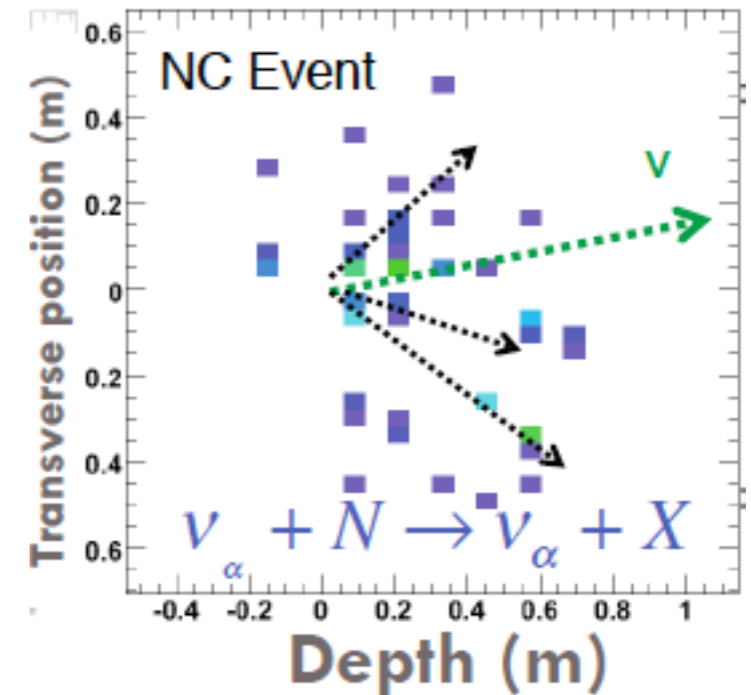


MINOS/T2K: $\delta_{CP}=0, \theta_{23}=\pi/4, \Delta m^2>0$

Sterile Neutrinos (?)

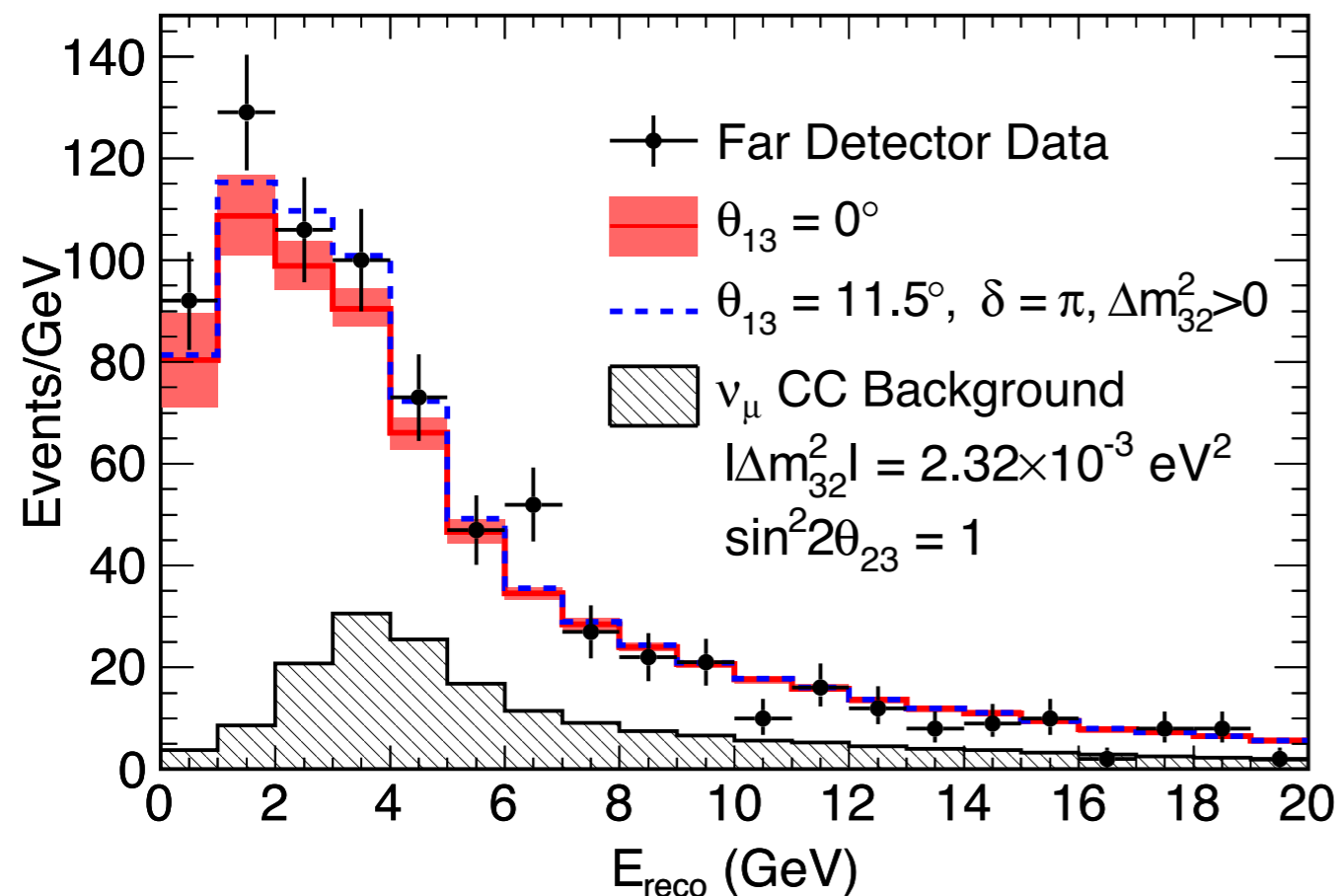
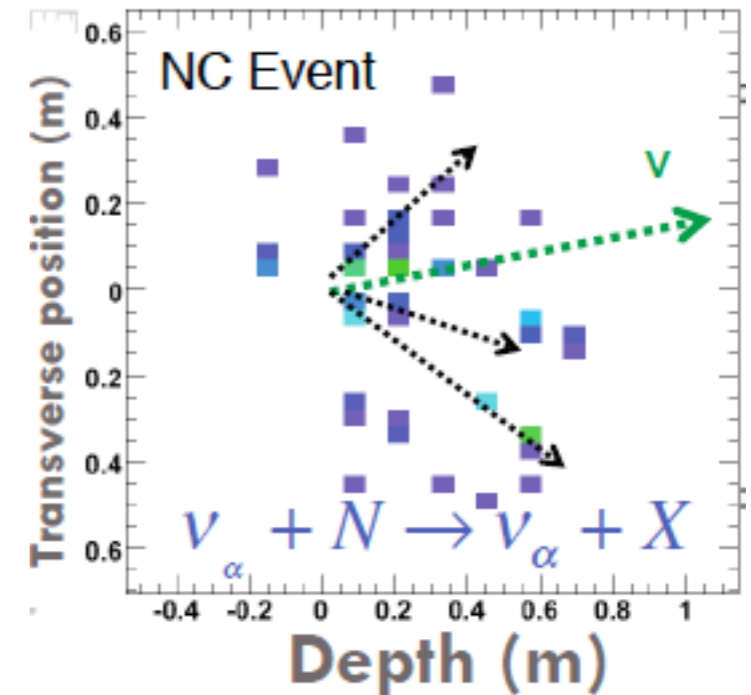
Sterile Neutrinos (?)

Relatively easy to identify: NO long track
Partial energy measurement (shower only):



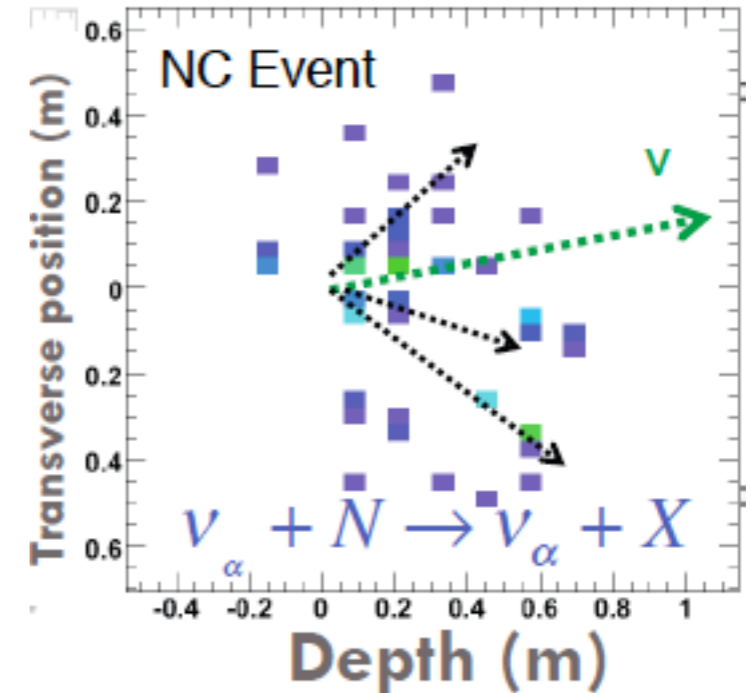
Sterile Neutrinos (?)

Relatively easy to identify: NO long track
Partial energy measurement (shower only):

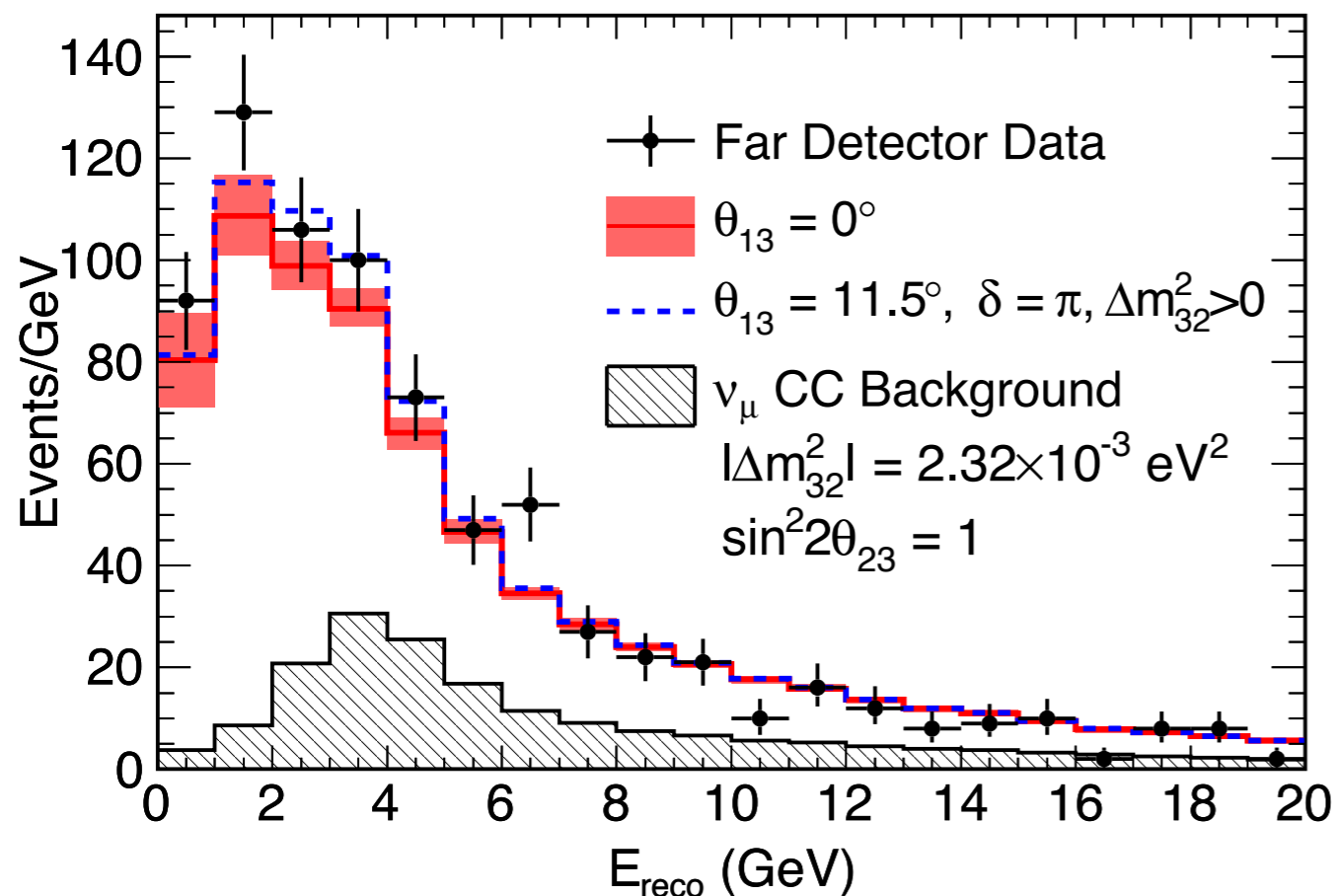


Sterile Neutrinos (?)

Relatively easy to identify: NO long track
 Partial energy measurement (shower only):



No evidence for sterile ν 's



Result

$$P(\nu_{\mu} \rightarrow \nu_s) / (1 - P(\nu_{\mu} \rightarrow \nu_{\mu})) \leq 0.22 (90\% \text{ CL})$$

Graphical Summary

Graphical Summary

$|\Delta m^2|$ now known to $\sim 5\%$

Graphical Summary

$|\Delta m^2|$ now known to $\sim 5\%$

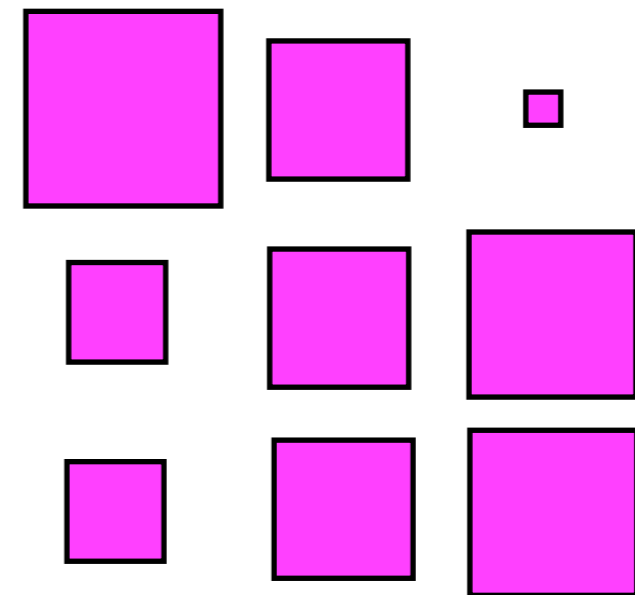
$|U_{\alpha i}|$ also known today to $\sim 5\%$
(except U_{e3})

Graphical Summary

$|\Delta m^2|$ now known to $\sim 5\%$

$|U_{\alpha i}|$ also known today to $\sim 5\%$
(except U_{e3})

Area of each square $\propto |U_{\alpha i}|^2$



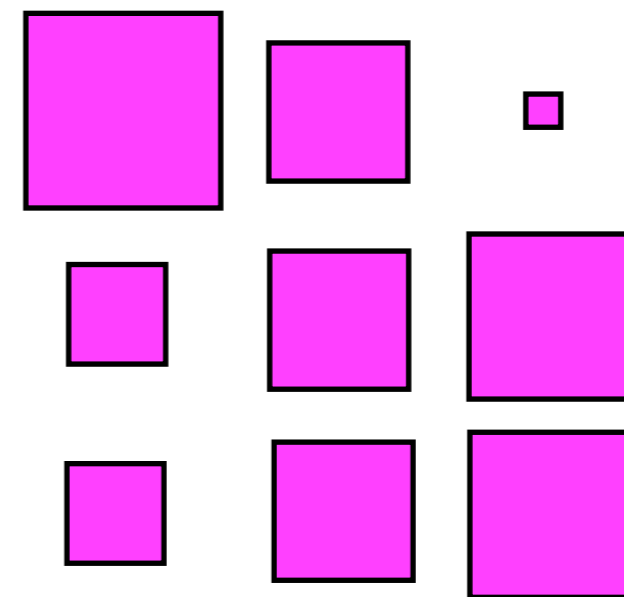
PMNS Matrix, graphically

Graphical Summary

$|\Delta m^2|$ now known to $\sim 5\%$

$|U_{\alpha i}|$ also known today to $\sim 5\%$
(except U_{e3})

Area of each square $\propto |U_{\alpha i}|^2$



PMNS Matrix, graphically

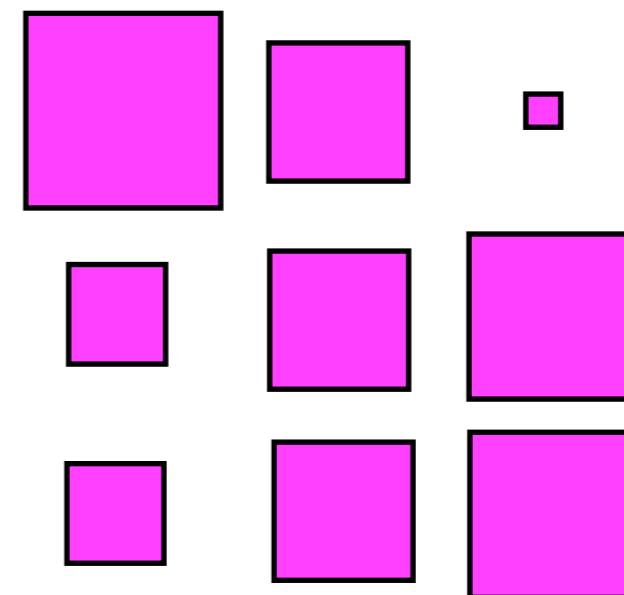


Graphical Summary

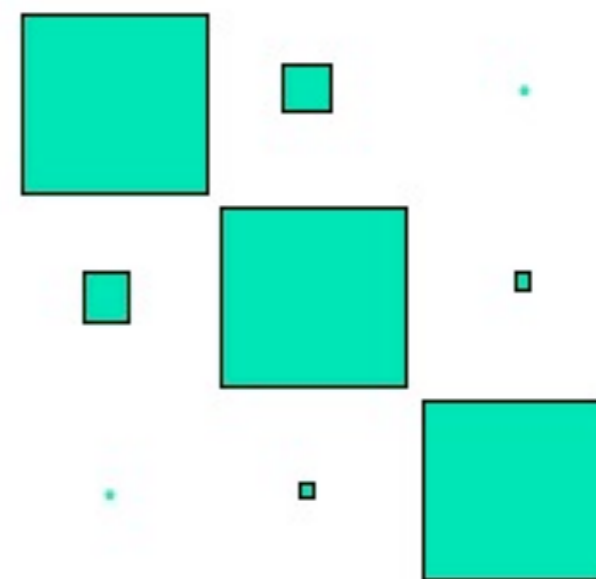
$|\Delta m^2|$ now known to $\sim 5\%$

$|U_{\alpha i}|$ also known today to $\sim 5\%$
(except U_{e3})

Area of each square $\propto |U_{\alpha i}|^2$



PMNS Matrix, graphically



CKM Matrix, graphically

Oscillations - Summary

- There has been great progress in the last decade and a half in determination of oscillation parameters
- But few (mass hierarchy, CP phase) remain to be determined
- Equally importantly is the possibility of surprises
- And then there are the really BIG questions:
 - Why is the mixing matrix the way it is?
 - Why is it so different from the quark sector?
 - Why are neutrino masses so small?
 - Are neutrinos responsible for us being here today?
- Too bad Pontecorvo is not around today

Velocity (Speed)

Velocity (Speed)

$$\text{Velocity} = \text{Distance/Time} = D / (t_{\text{stop}} - t_{\text{start}})$$

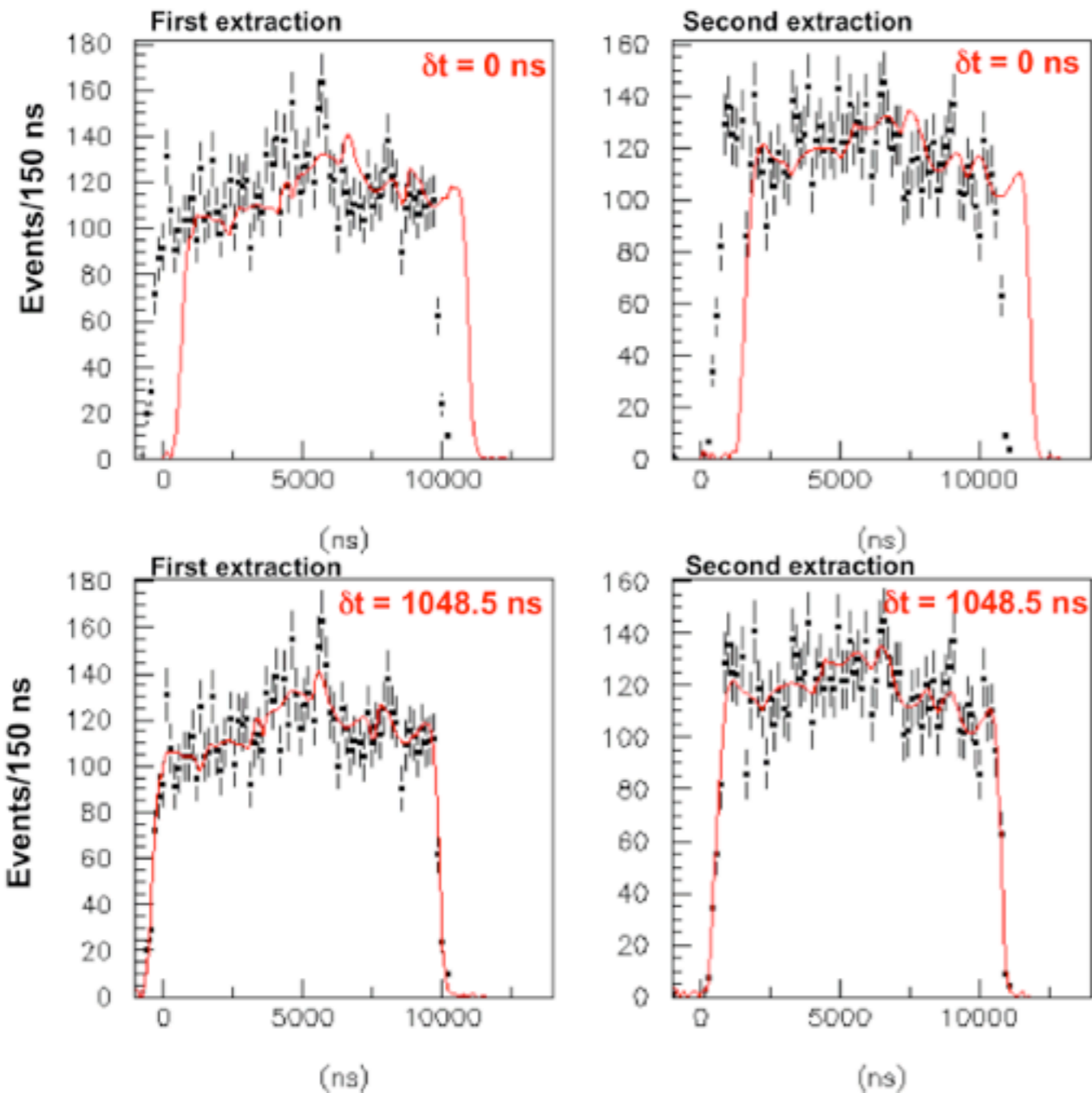


t_{start}



t_{stop}

OPERA Results - 1st Run



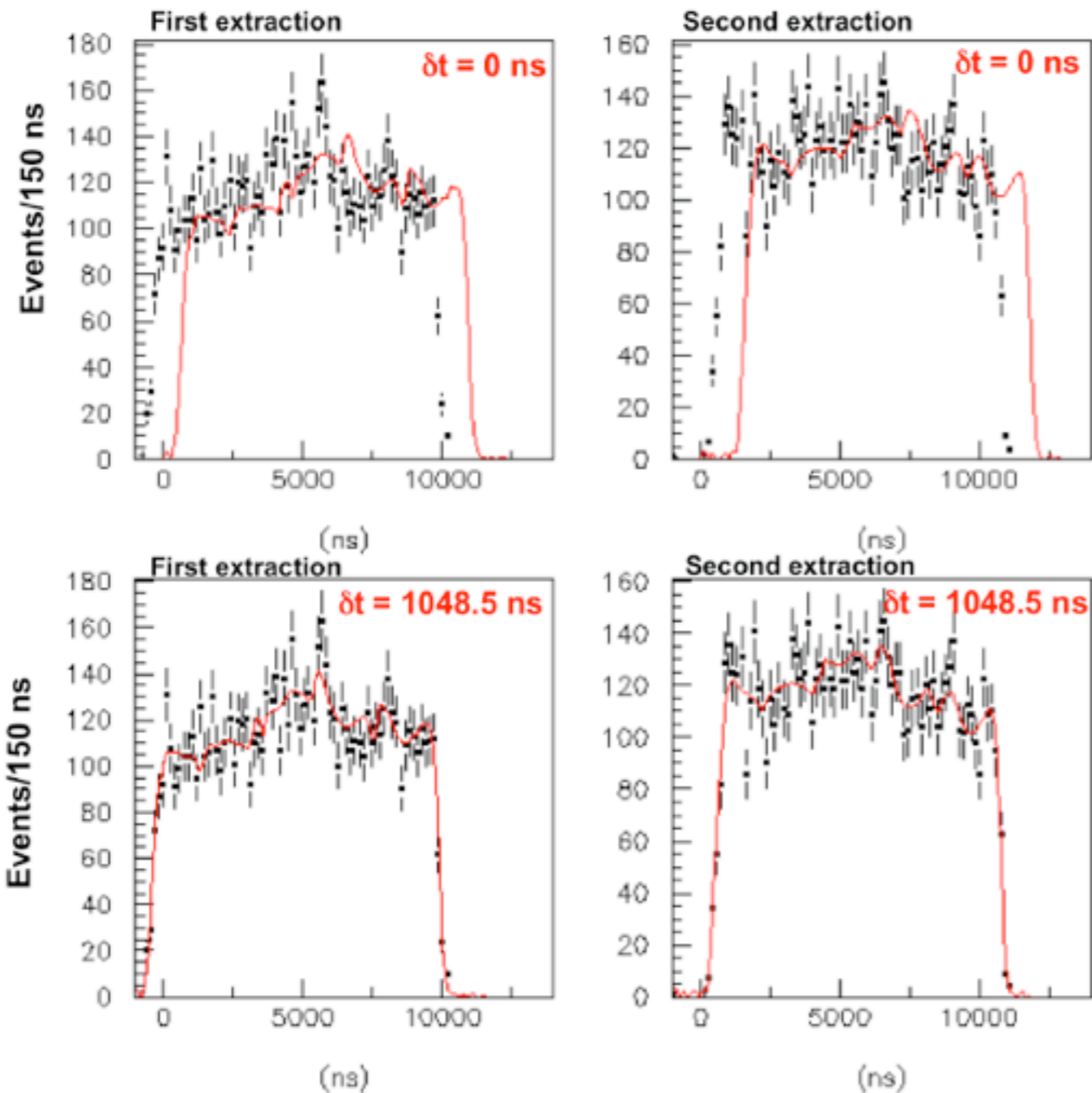
— Proton beam time structure
● ν event time distribution

Result

$$\delta t = (57.8 \pm 7.8(\text{stat.}) -5.9^{+8.3}(\text{sys.})) \text{ ns}$$

$$(\nu - c)/c = (2.37 \pm 0.32(\text{stat.}) -0.24^{+0.34}(\text{sys.})) \times 10^{-5}$$

OPERA Results - 1st Run



7235 internal events (CC and NC)
7988 external events (CC)

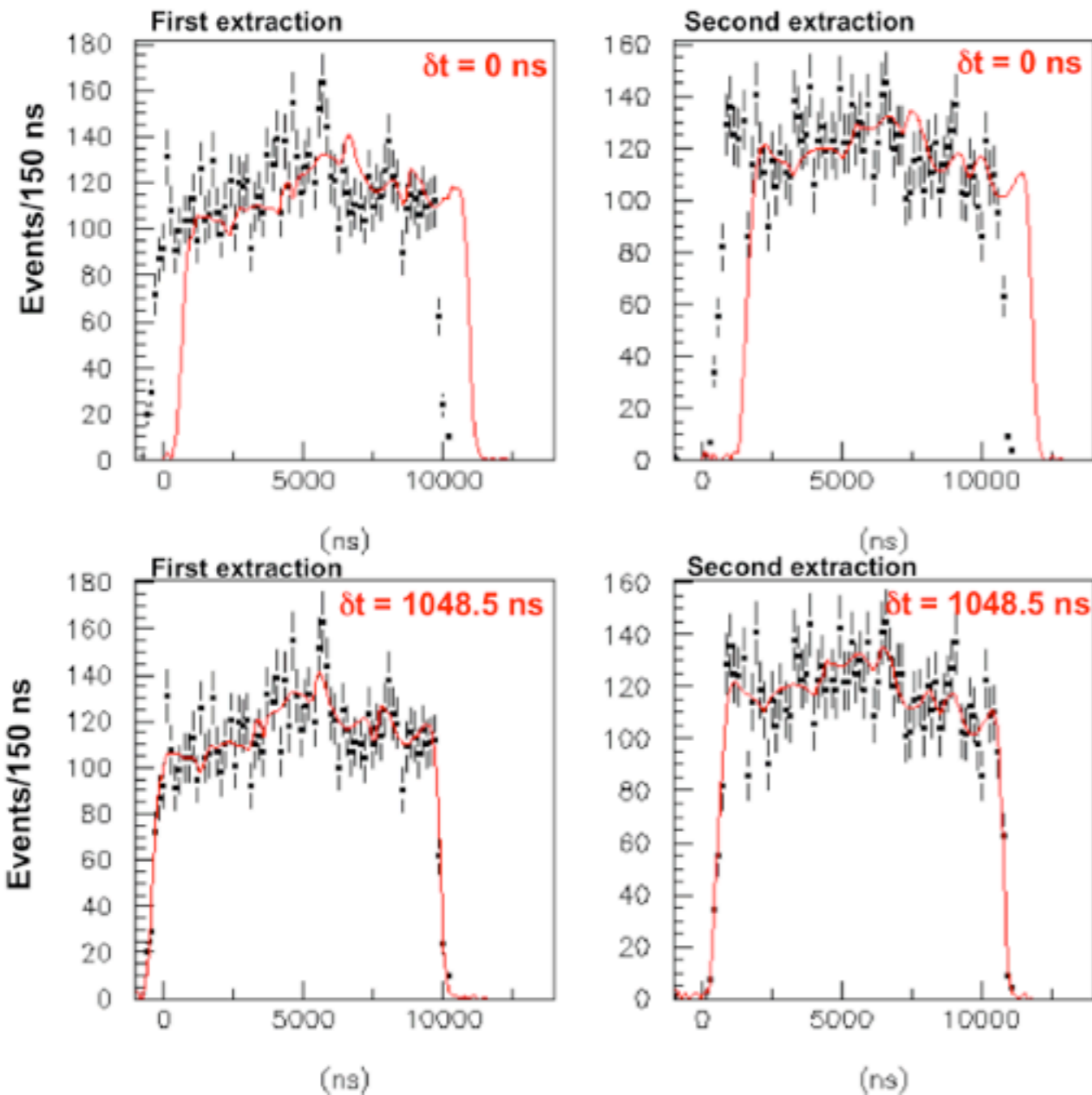
— Proton beam time structure
● ν event time distribution

Result

$$\delta t = (57.8 \pm 7.8(\text{stat.}) -5.9^{+8.3}(\text{sys.})) \text{ ns}$$

$$(\nu - c)/c = (2.37 \pm 0.32(\text{stat.}) -0.24^{+0.34}(\text{sys.})) \times 10^{-5}$$

OPERA Results - 1st Run



7235 internal events (CC and NC)
7988 external events (CC)

— Proton beam time structure
● ν event time distribution

Clearly the most useful information comes from neutrino events at either the leading or trailing edges

Result

$$\delta t = (57.8 \pm 7.8(\text{stat.}) -5.9^{+8.3}(\text{sys.})) \text{ ns}$$

$$(\nu - c)/c = (2.37 \pm 0.32(\text{stat.}) -0.24^{+0.34}(\text{sys.})) \times 10^{-5}$$

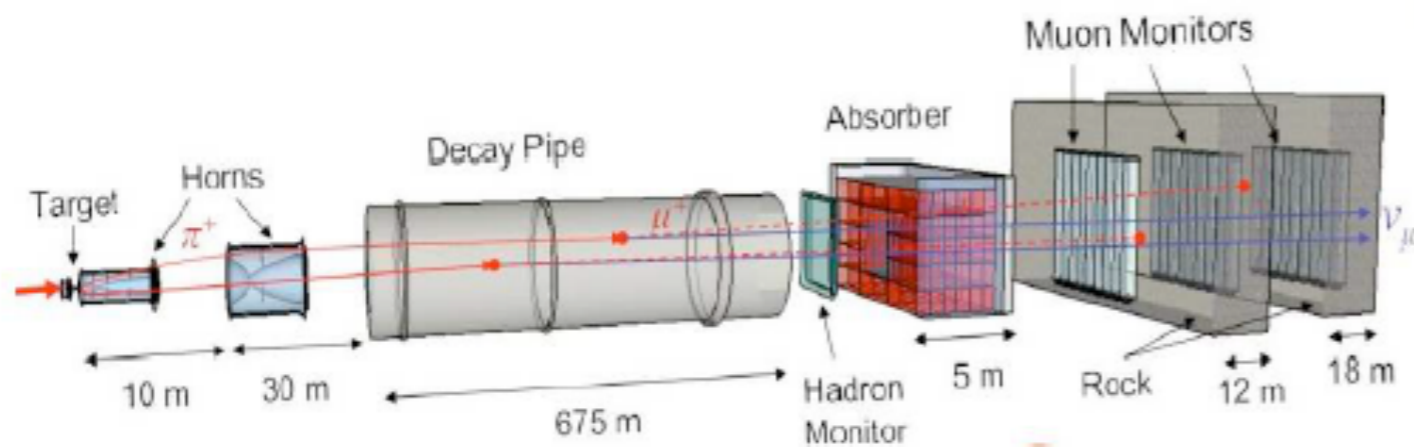
OPERA / MINOS Comparison

OPERA / MINOS Comparison

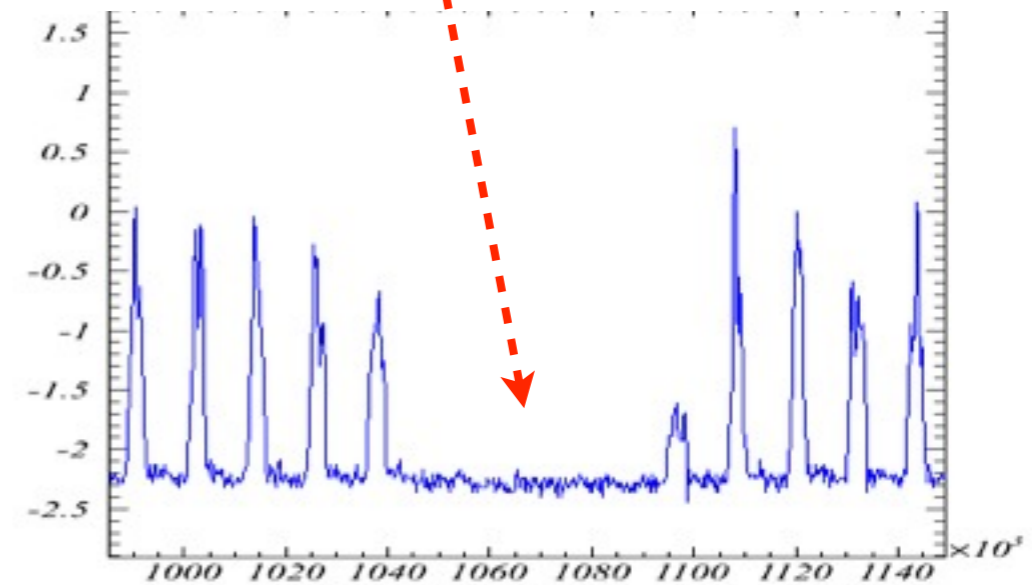
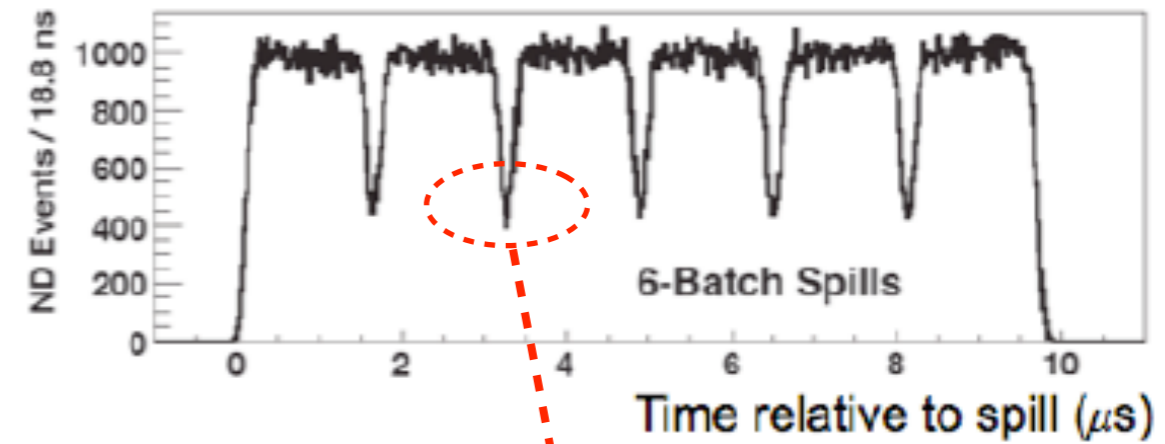
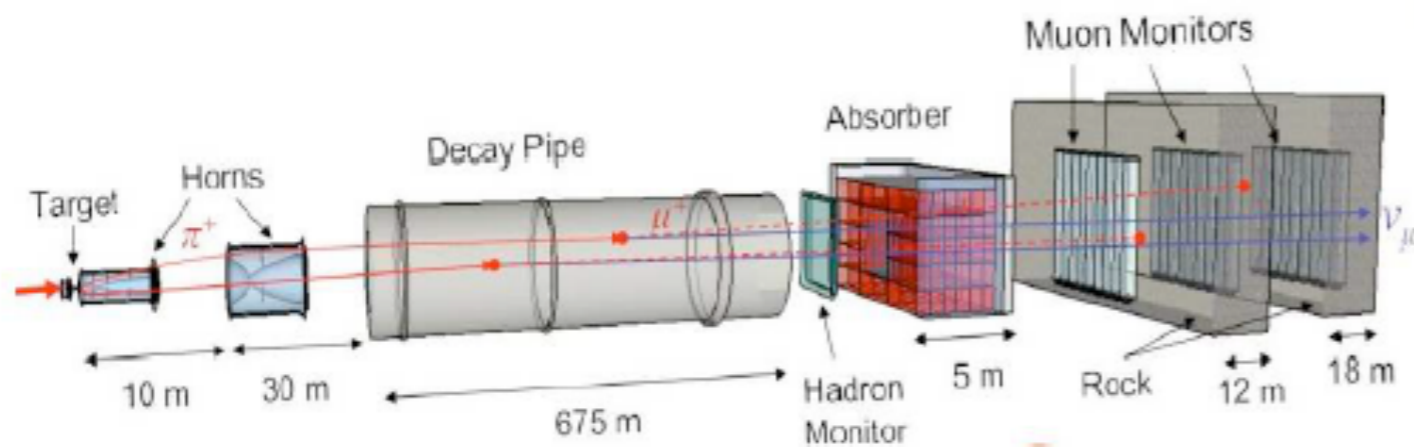
Parameter	OPERA	MINOS
Location	CERN->Gran Sasso	Fermilab->Soudan
Distance	731 km	735 km
Main Physics Goal	ν_{τ} appearance	ν_{μ} disappearance
Configuration	No Near Detector	Near Detector
Det. composition	Emulsion/Pb	Scintillator/Fe
ν Energy	High, <17 GeV>	Low, <3 GeV>
Data taking	2009-2011	2005-2006, 2011
Beam structure	2 10.5 μ s, 200 MHz	6 1.6 μ s, 53 MHz
Access to FD	Highway (10 km)	Elevator (2000 ft)

MINOS Beam

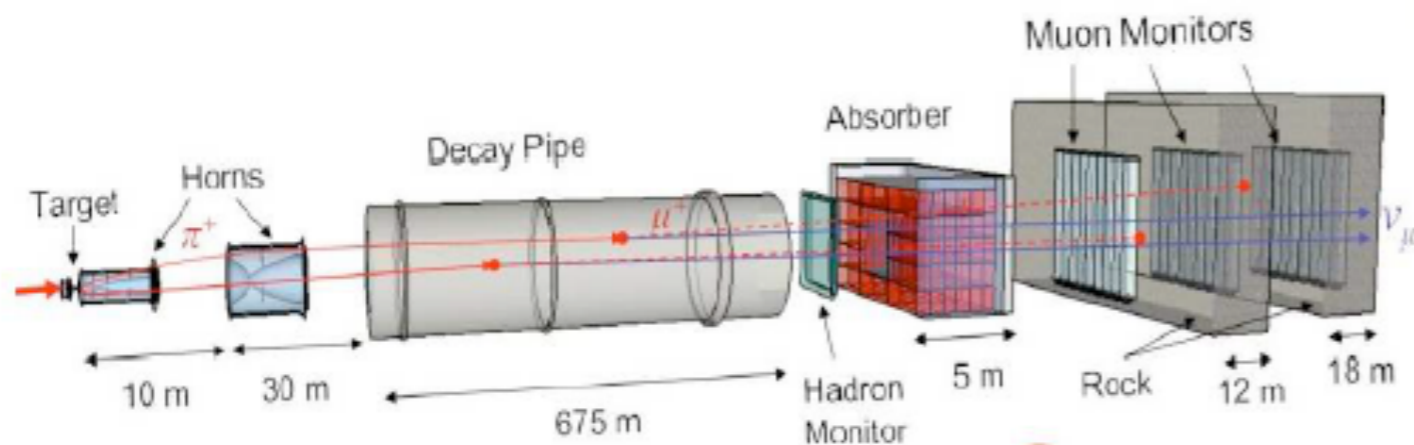
MINOS Beam



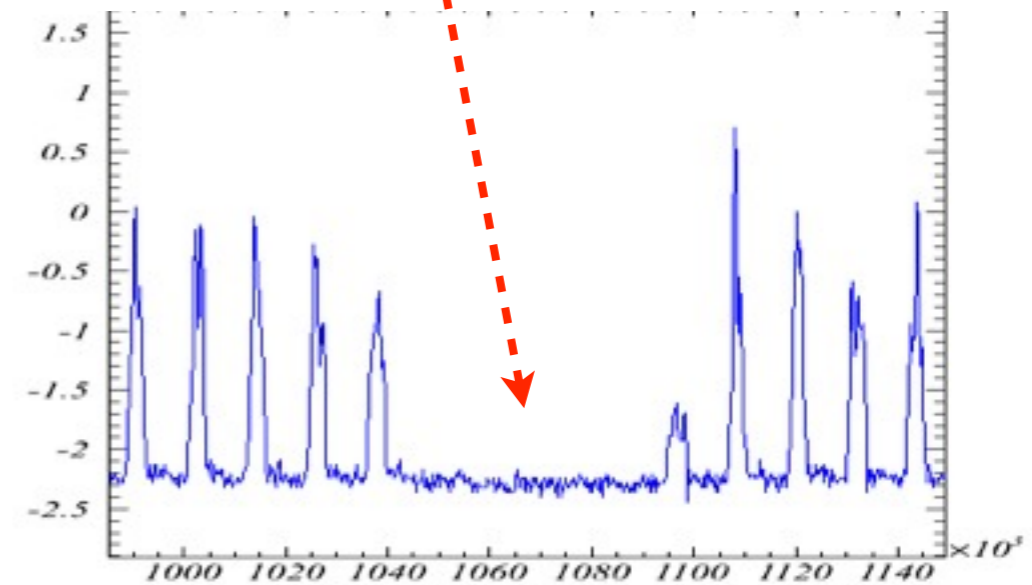
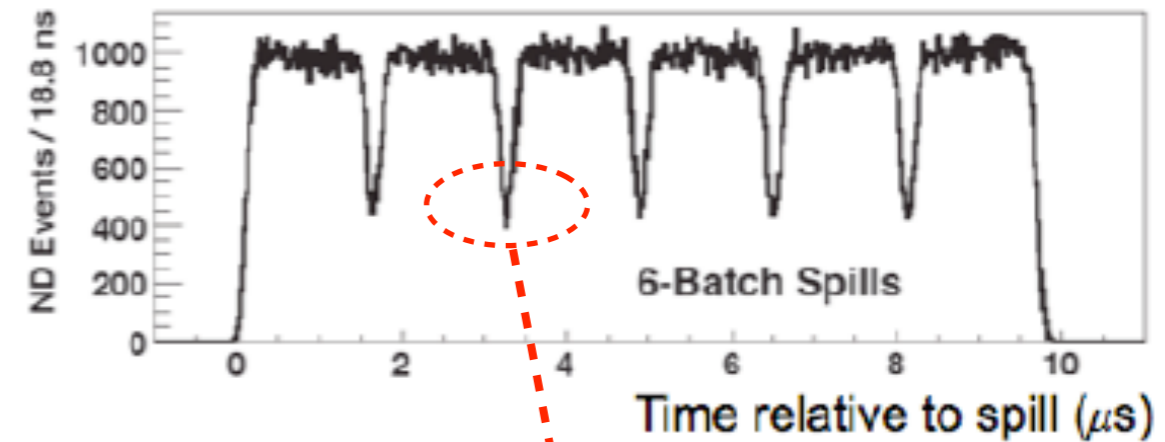
MINOS Beam



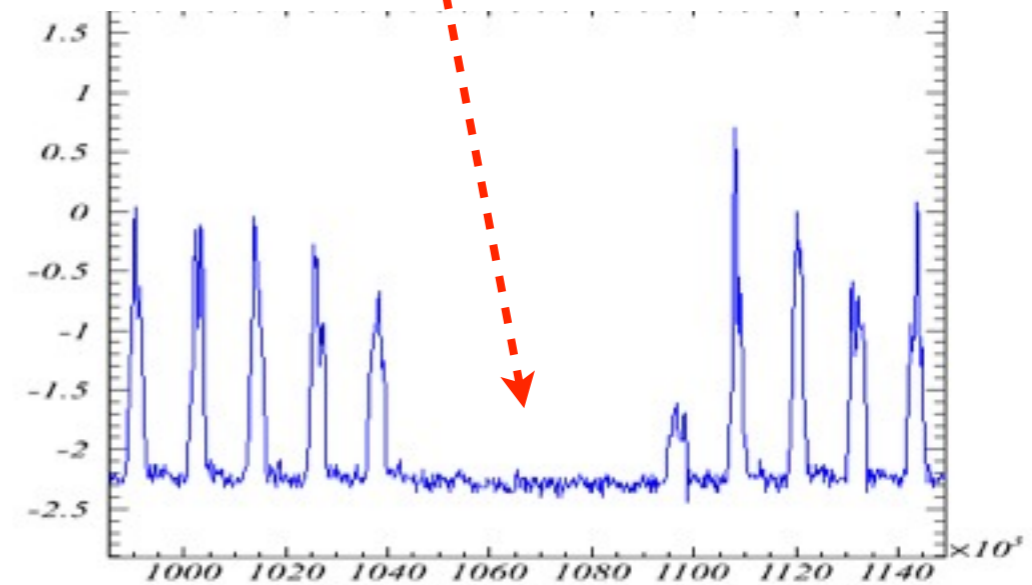
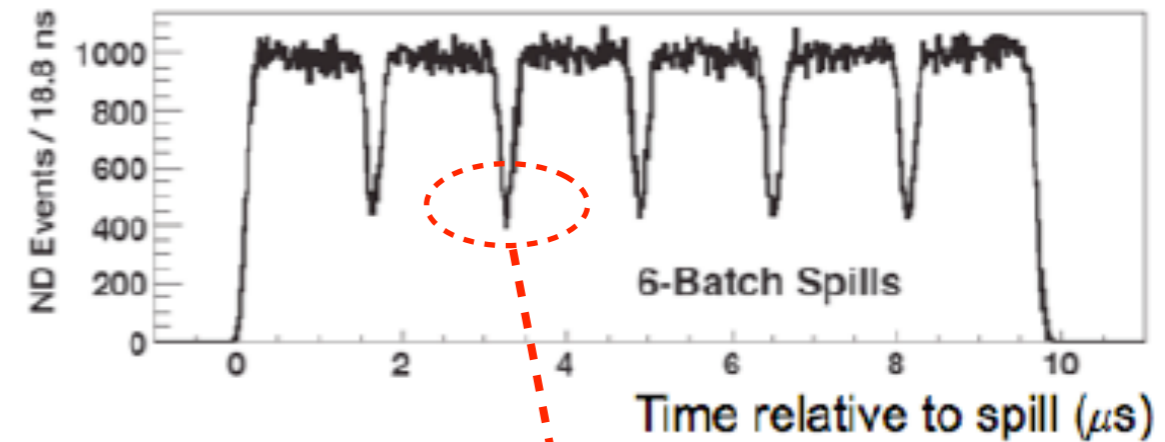
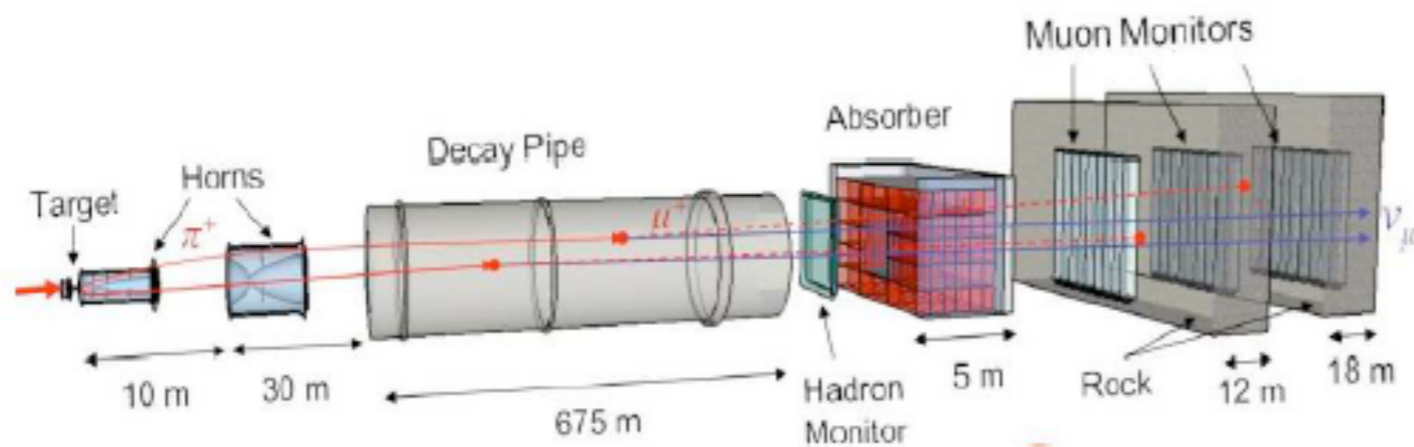
MINOS Beam



- ~ 1.8 s cycle

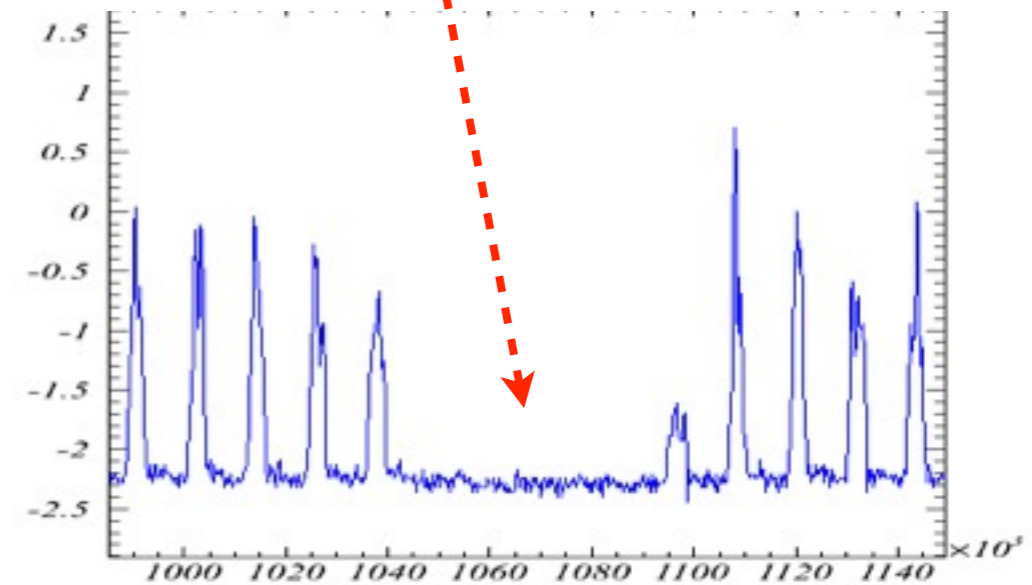
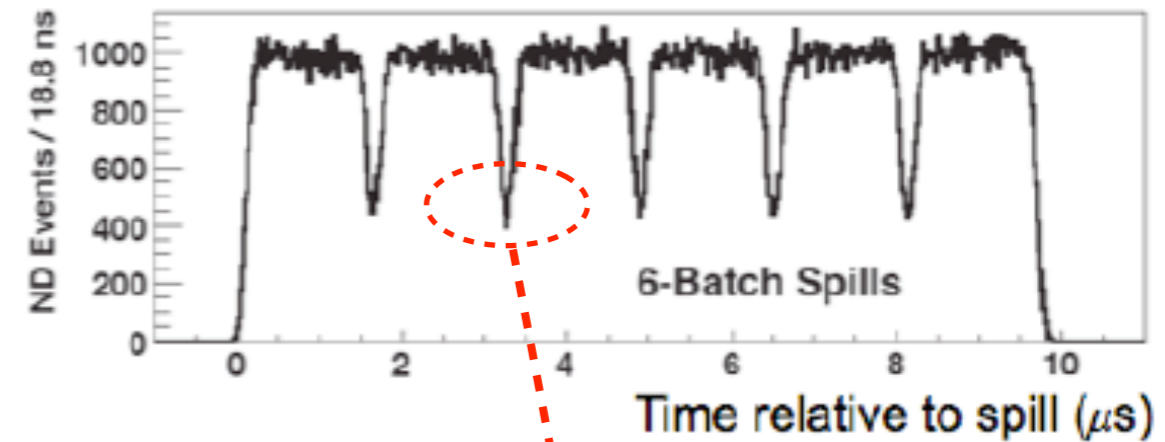
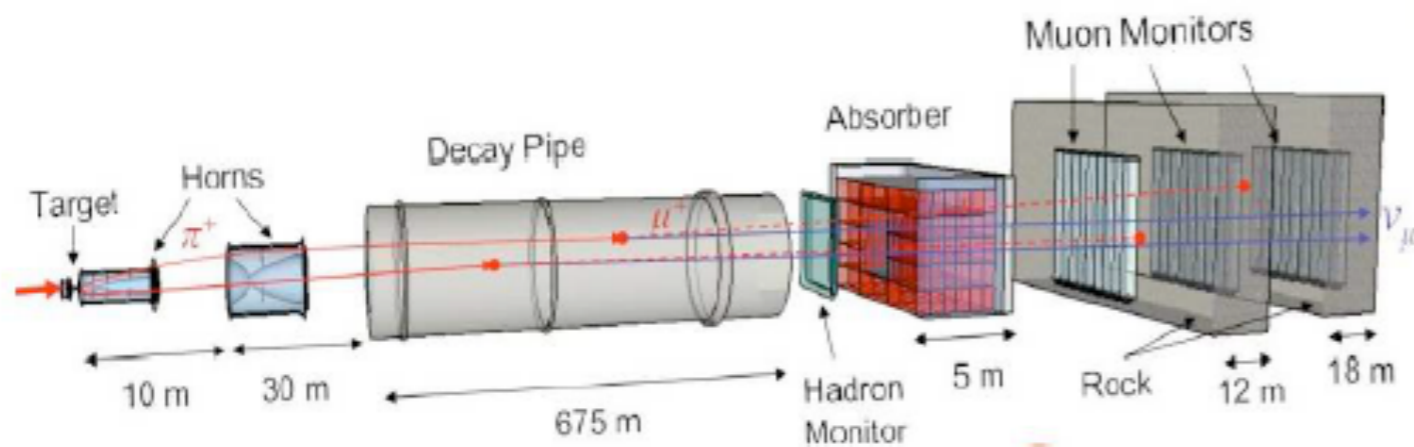


MINOS Beam



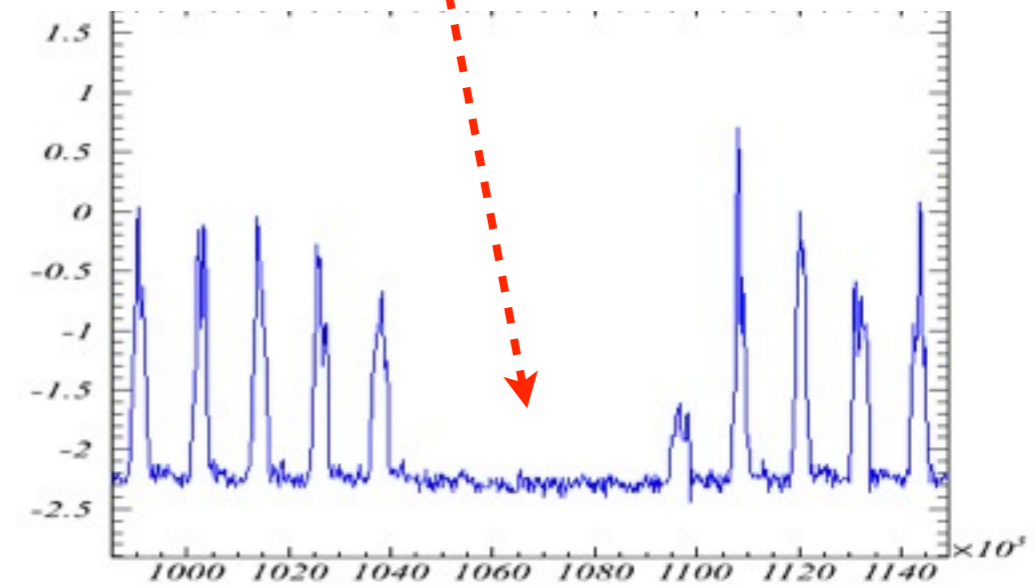
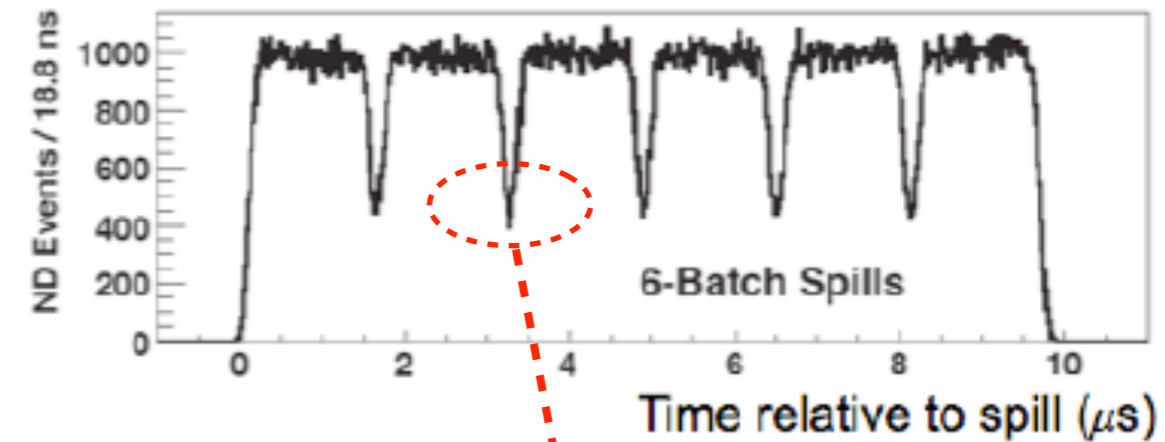
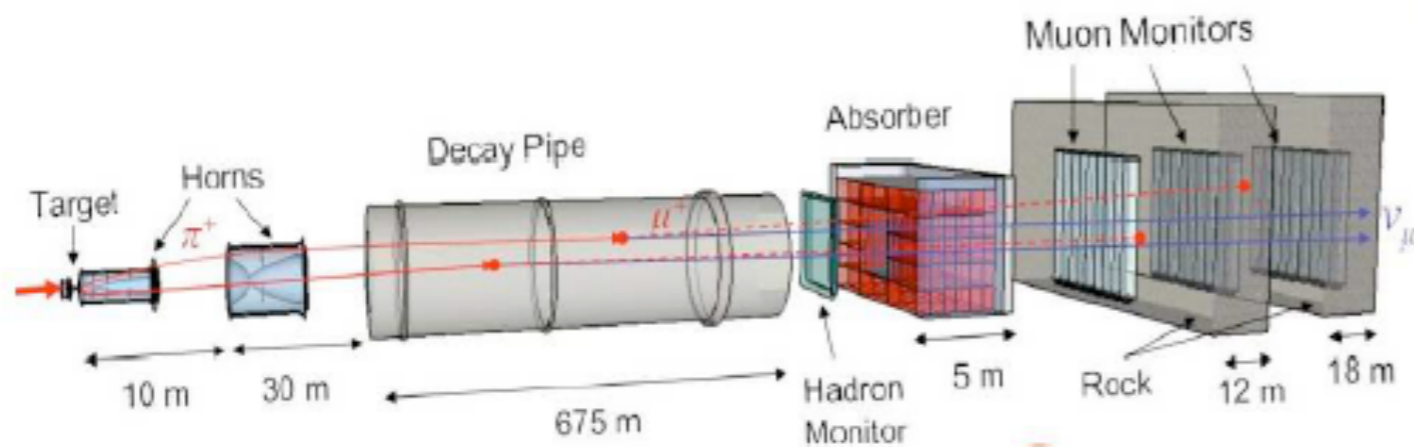
- ~ 1.8 s cycle
- Beam spill 9.6 μs long

MINOS Beam



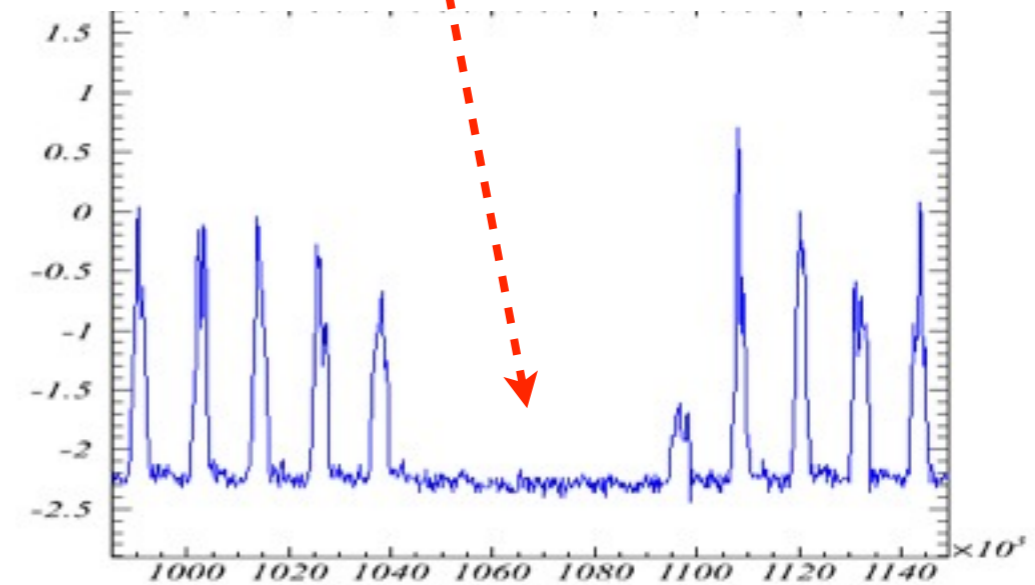
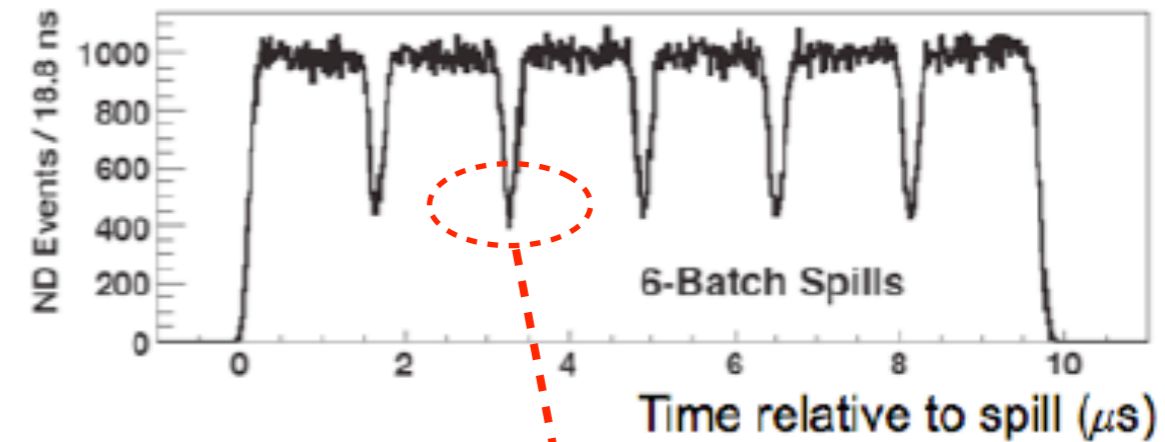
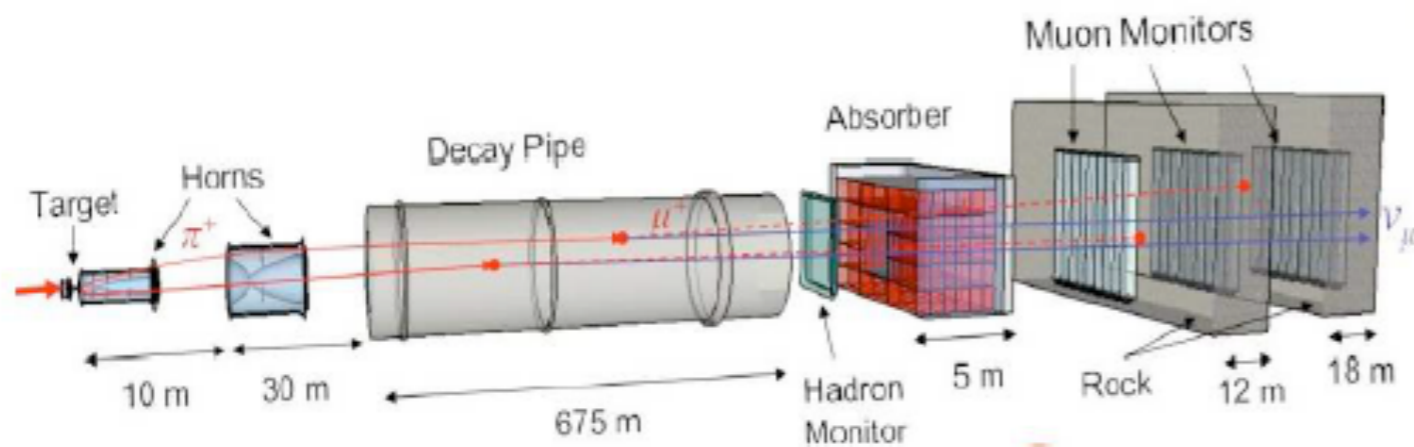
- ~ 1.8 s cycle
- Beam spill 9.6 μs long
- 5 or 6 batches in spill

MINOS Beam



- ~ 1.8 s cycle
- Beam spill $9.6 \mu\text{s}$ long
- 5 or 6 batches in spill
- Each batch $1.52 \mu\text{s}$ long, separated by 90 ns

MINOS Beam



- ~ 1.8 s cycle
- Beam spill $9.6 \mu\text{s}$ long
- 5 or 6 batches in spill
- Each batch $1.52 \mu\text{s}$ long, separated by 90 ns
- Modulated by 53 MHz (18.8 ns)

3 Measurement Phases

3 Measurement Phases

I. Analysis of data already taken

3 Measurement Phases

- I. Analysis of data already taken
 - Much higher statistics (x8)

3 Measurement Phases

I. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)

3 Measurement Phases

I. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis

3 Measurement Phases

I. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, $\sim 2-3$ months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3. 2013 data

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3. 2013 data

- Data will be taken at higher energy

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, $\sim 2-3$ months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3. 2013 data

- Data will be taken at higher energy
- Build on experience from initial 2 phases

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3. 2013 data

- Data will be taken at higher energy
- Build on experience from initial 2 phases
- Aim for single RF bucket resolution

3 Measurement Phases

1. Analysis of data already taken

- Much higher statistics (x8)
- Reduced systematics (Delays, GPS, relative timing)
- Improved analysis
- Aim for ~ 15 ns uncertainty, ~ 2 -3 months

2. New data with improved timing hardware

- 2012 data before shutdown (scheduled for May 1)
- Timing at 3 locations: extraction (beam wall current), ND, and FD
- Redundant timing
- Aim for < 10 ns, 5-8 months

3. 2013 data

- Data will be taken at higher energy
- Build on experience from initial 2 phases
- Aim for single RF bucket resolution
- Remeasurement of distance (?)

MINOS Beam

Modified J.Thomas slide

MINOS Beam



Modified J.Thomas slide

MINOS Beam



t'_{start} measurement from wall current monitor

Modified J.Thomas slide

MINOS Beam



protons highly relativistic (120 GeV/c) $\Delta t < 100$ ps

t'_{start} measurement from wall current monitor

Modified J.Thomas slide

MINOS Beam



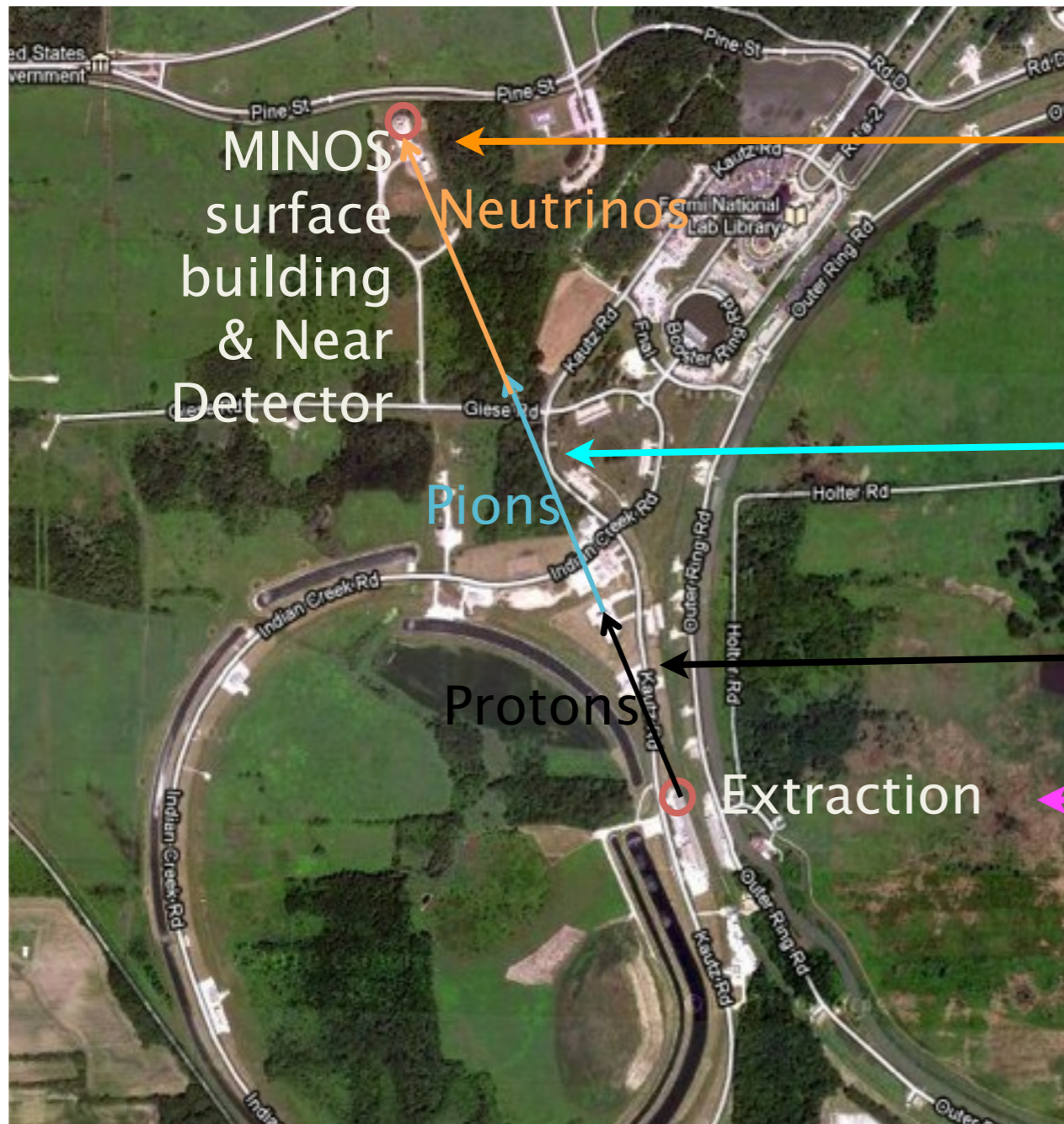
pions highly relativistic, ~ 10 GeV/c,
decay point not very relevant,
 $\Delta t \sim$ few hundred ps, calculable

protons highly relativistic (120
GeV/c) $\Delta t < 100$ ps

t'_{start} measurement from wall
current monitor

Modified J.Thomas slide

MINOS Beam



t_{start} measurement from neutrinos in the Near Detector

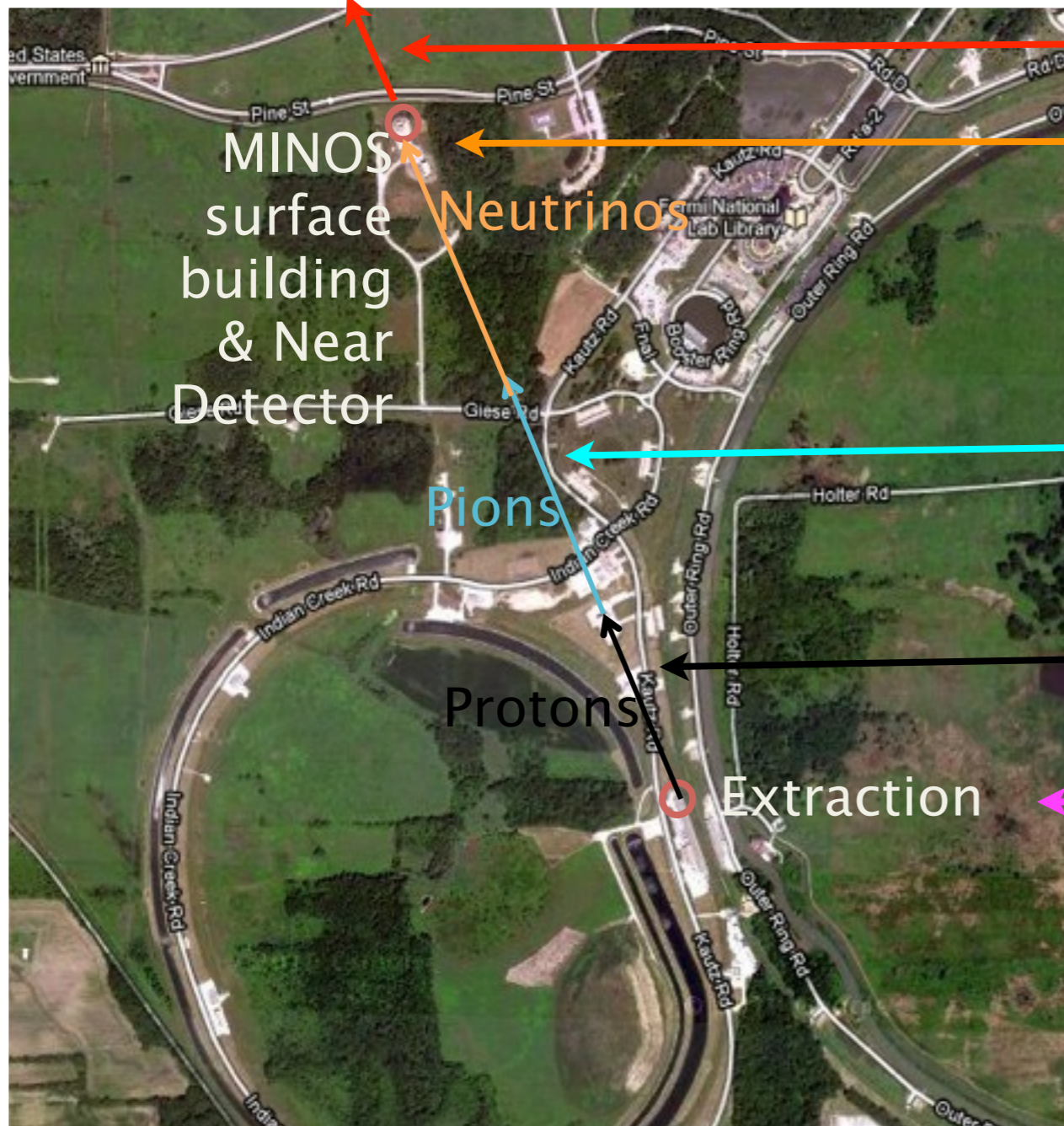
pions highly relativistic, ~ 10 GeV/c, decay point not very relevant, $\Delta t \sim$ few hundred ps, calculable

protons highly relativistic (120 GeV/c) $\Delta t < 100$ ps

t'_{start} measurement from wall current monitor

Modified J.Thomas slide

MINOS Beam



Neutrinos to Soudan (t_{stop})
 t_{start} measurement from
neutrinos in the Near Detector

pions highly relativistic, ~ 10 GeV/c,
decay point not very relevant,
 $\Delta t \sim$ few hundred ps, calculable

protons highly relativistic (120
GeV/c) $\Delta t < 100$ ps

t'_{start} measurement from wall
current monitor

Modified J.Thomas slide

Final Comments re TOF

- Superluminal neutrinos would drastically alter our picture of physics
- Even though no problem has been found in the OPERA analysis, physics community remains sceptical
- An independent experiment is required to test the results of OPERA
- Personally I am sceptical - it was NOT suggested as a neutrino property by Pontecorvo

Acknowledgements

- In my professional career I have been very fortunate in having a chance to collaborate with many younger, very gifted and very hard working colleagues
- Any success that I might have had is very much due to these collaborative efforts, both today and in the past
- I want to thank all my present and past collaborators for their many contributions